

Heavy ion collisions as a tool: which physics can be explored?

The XVth Quark confinement and the Hadron spectrum
conference
in Stavanger

02/08/2022

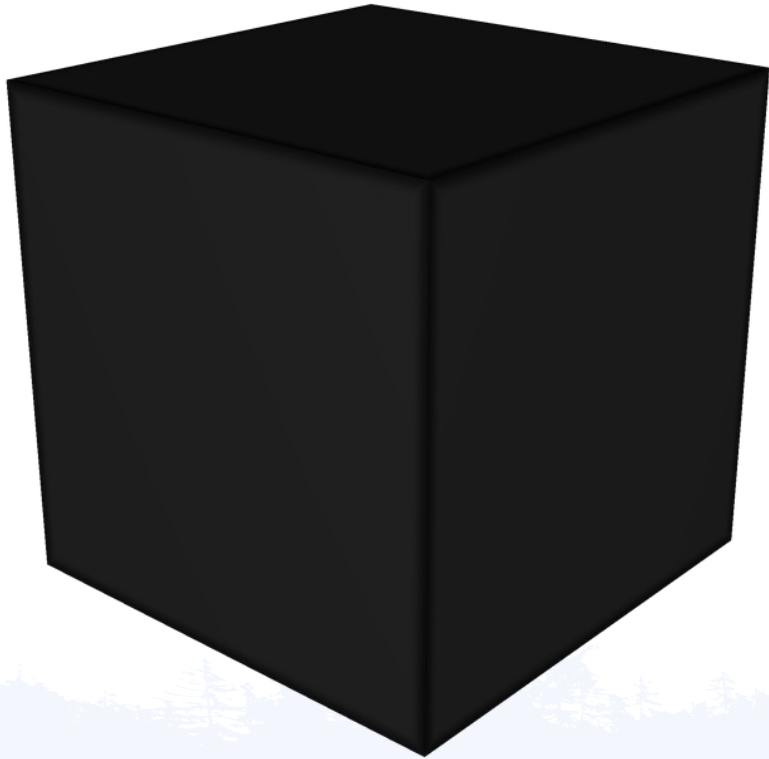
Nicolò Jacazio (Bologna University)

<https://about.me/jacazio>



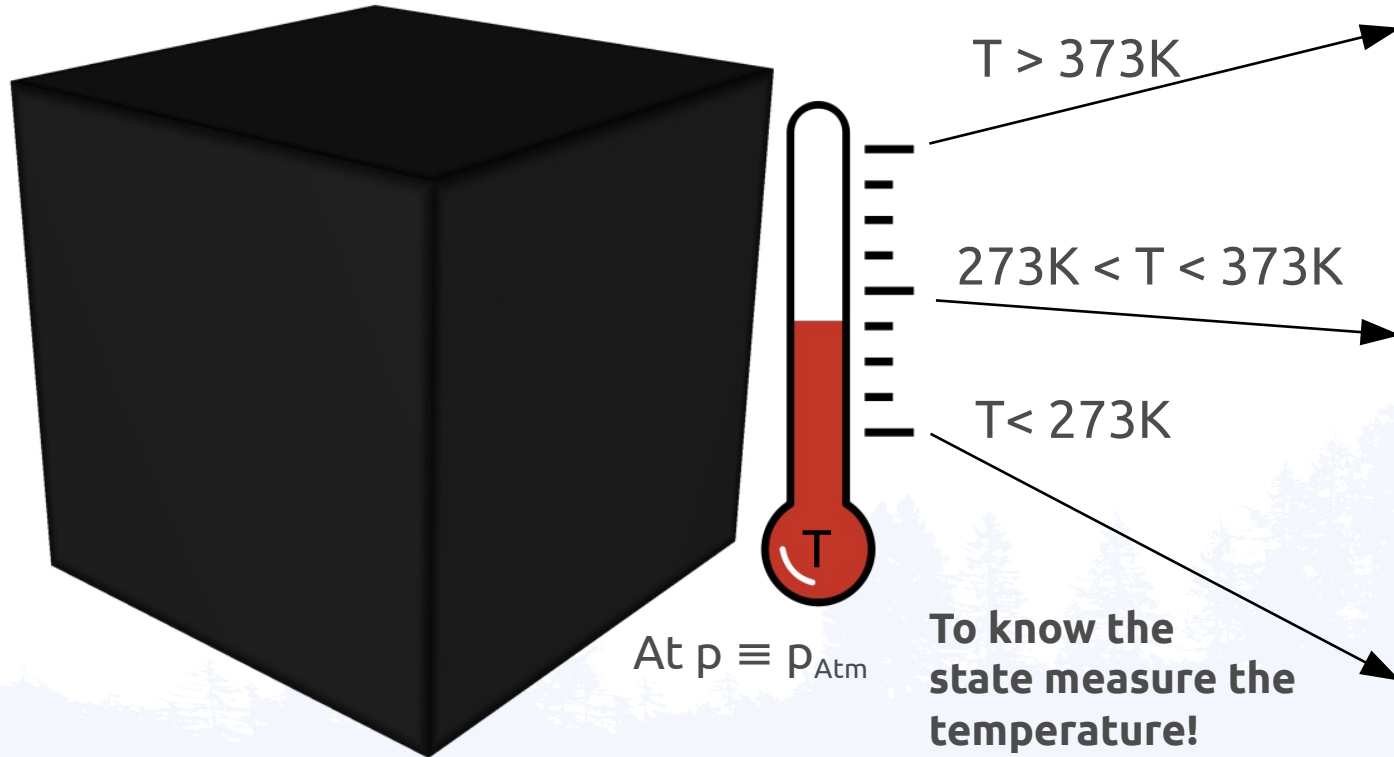
What's in the box?

- Imagine having a black box with water inside
- How to know what's really inside?



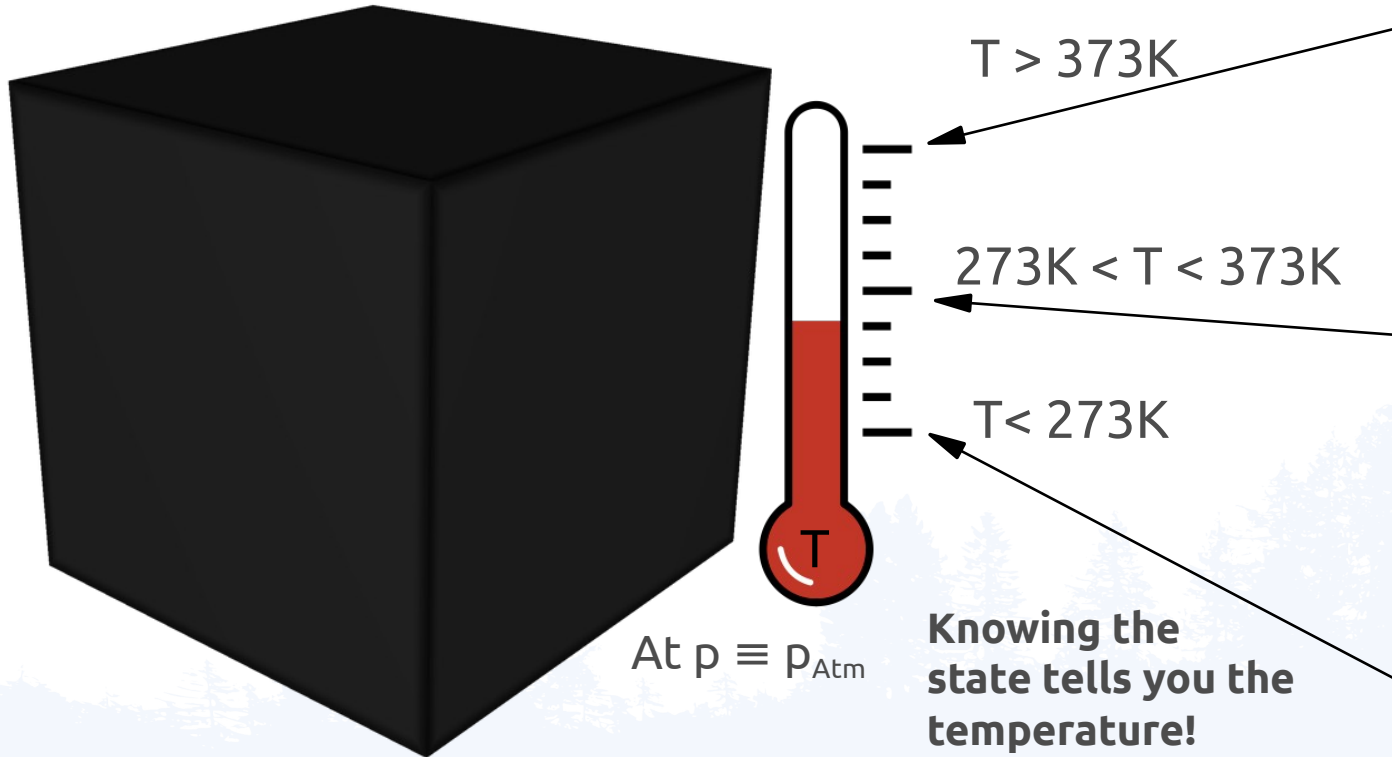
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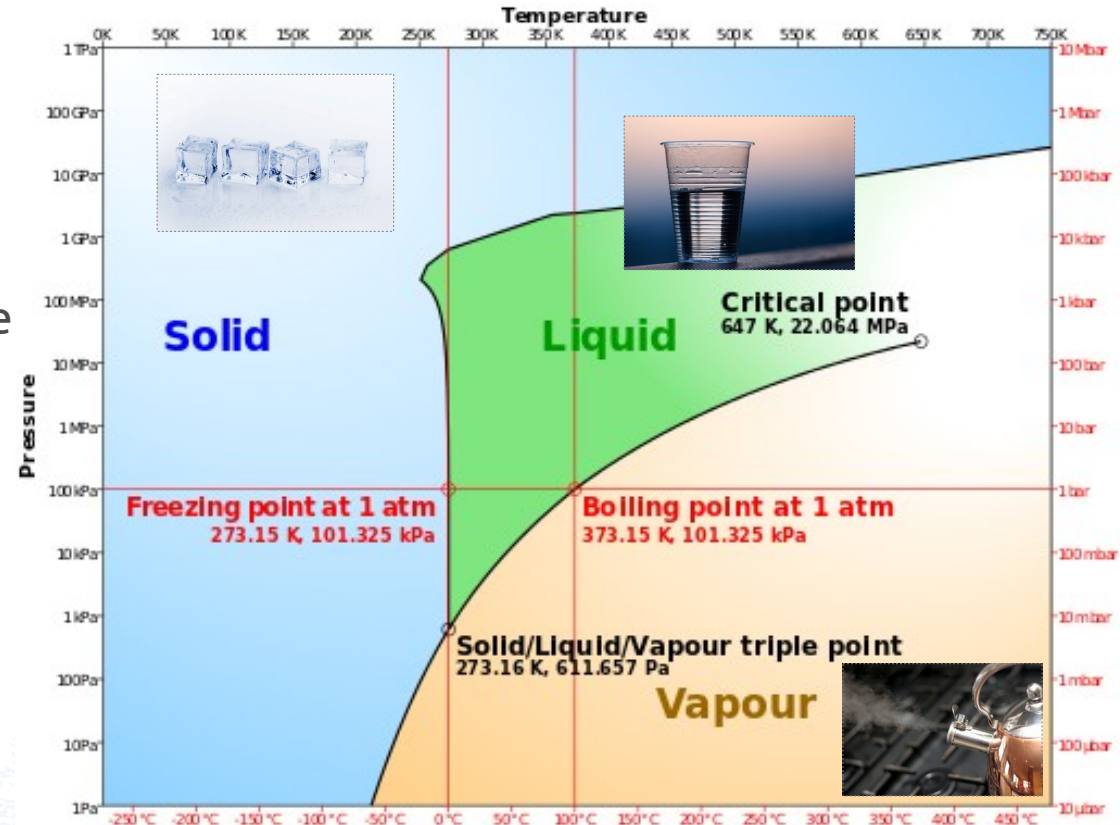


Knowing the state tells you the temperature!



Water phase diagram

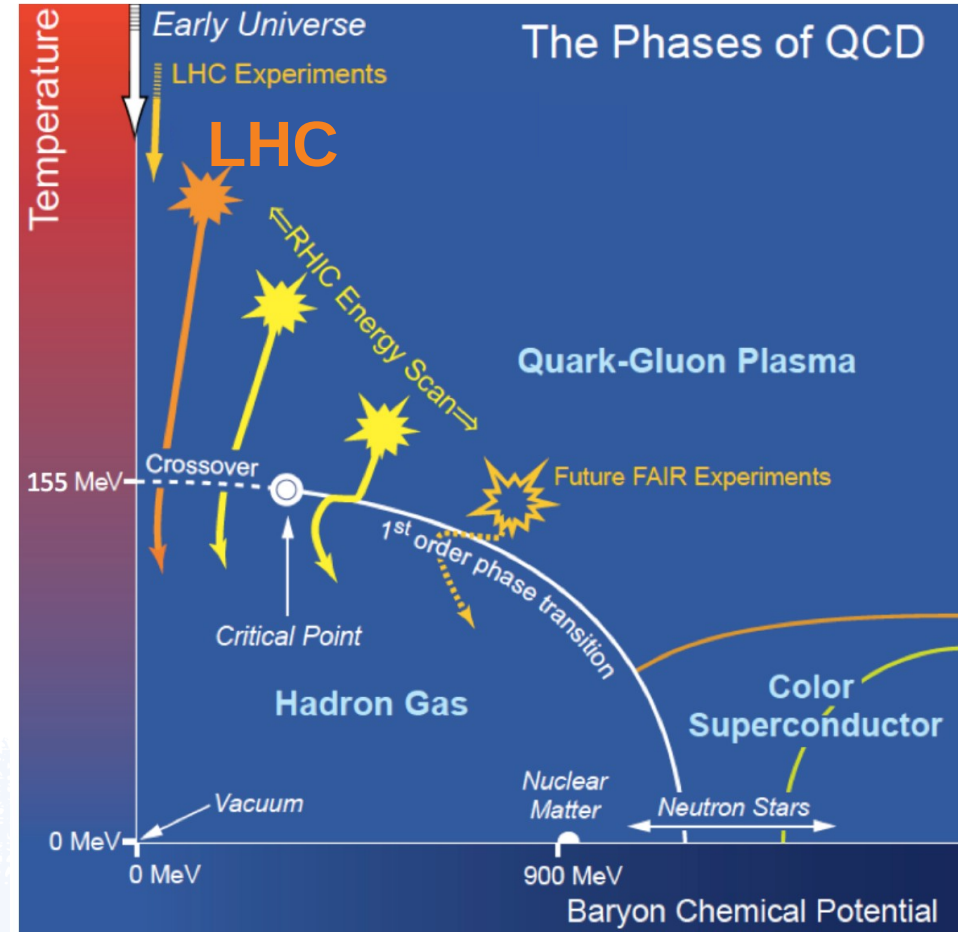
- The phase of the matter in the box depends on its thermodynamic state
- Three phases (at least) are available in the case of water
- Phase diagram investigated by varying pressure and temperature



QCD phase diagram

- The QCD phase diagram can be investigated in the same way
- High temperatures and low net baryon density are reached in heavy-ion collisions at the LHC
- Deconfined (strongly interacting) quark and gluons → the **Quark-Gluon Plasma**
- Partons moving over distances larger than the typical size of hadrons

Main physics goal in heavy ion physics!

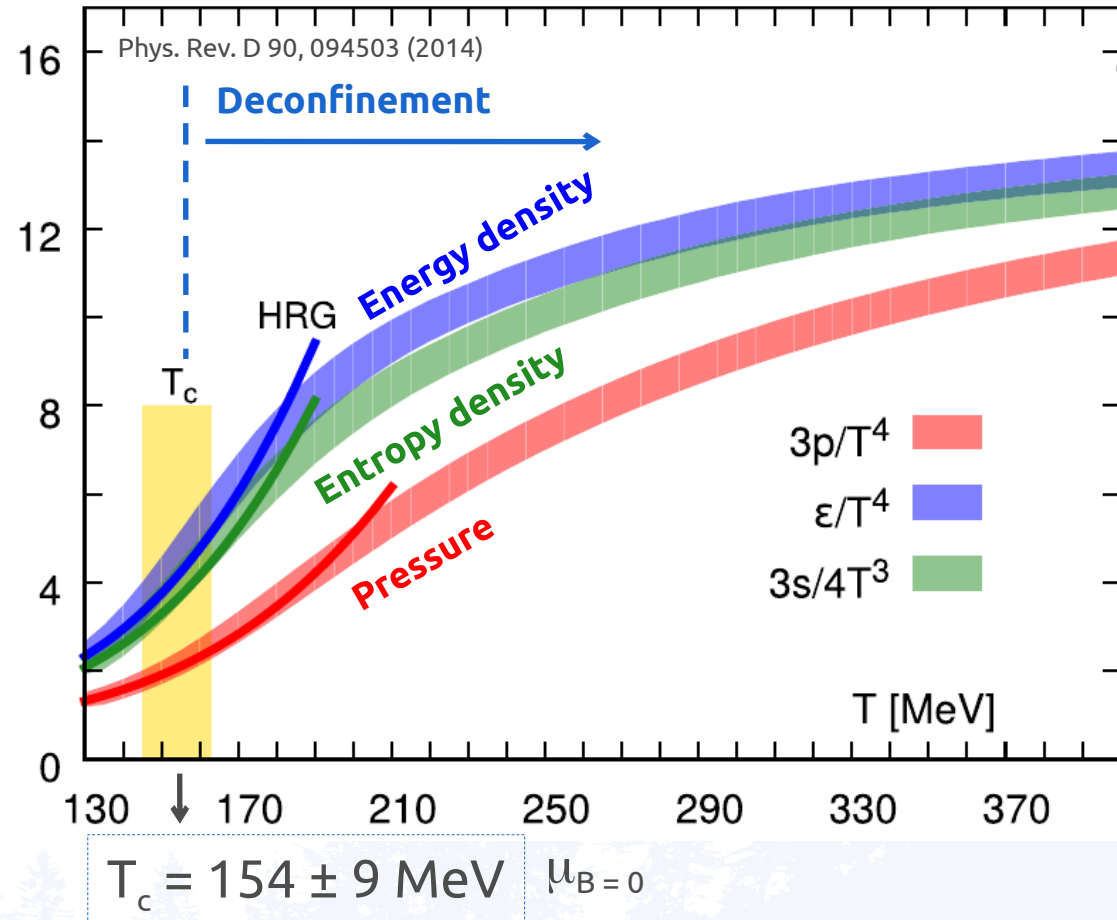


Heavy-ion collisions

Computations from lattice QCD identify a critical temperature T_c

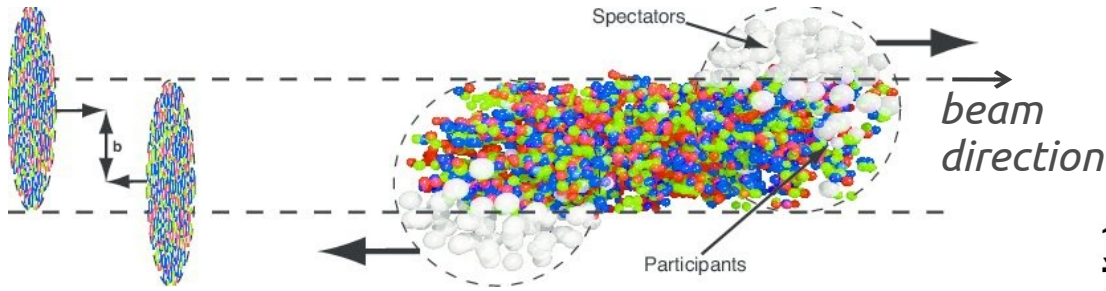
- $T > T_c$: deconfined phase, quarks and gluons are the degrees of freedom
- $T < T_c$: confined phase, hadrons are the degrees of freedom

Systems with energy densities $> 1 \text{ GeV/fm}^3$ produce a deconfined medium

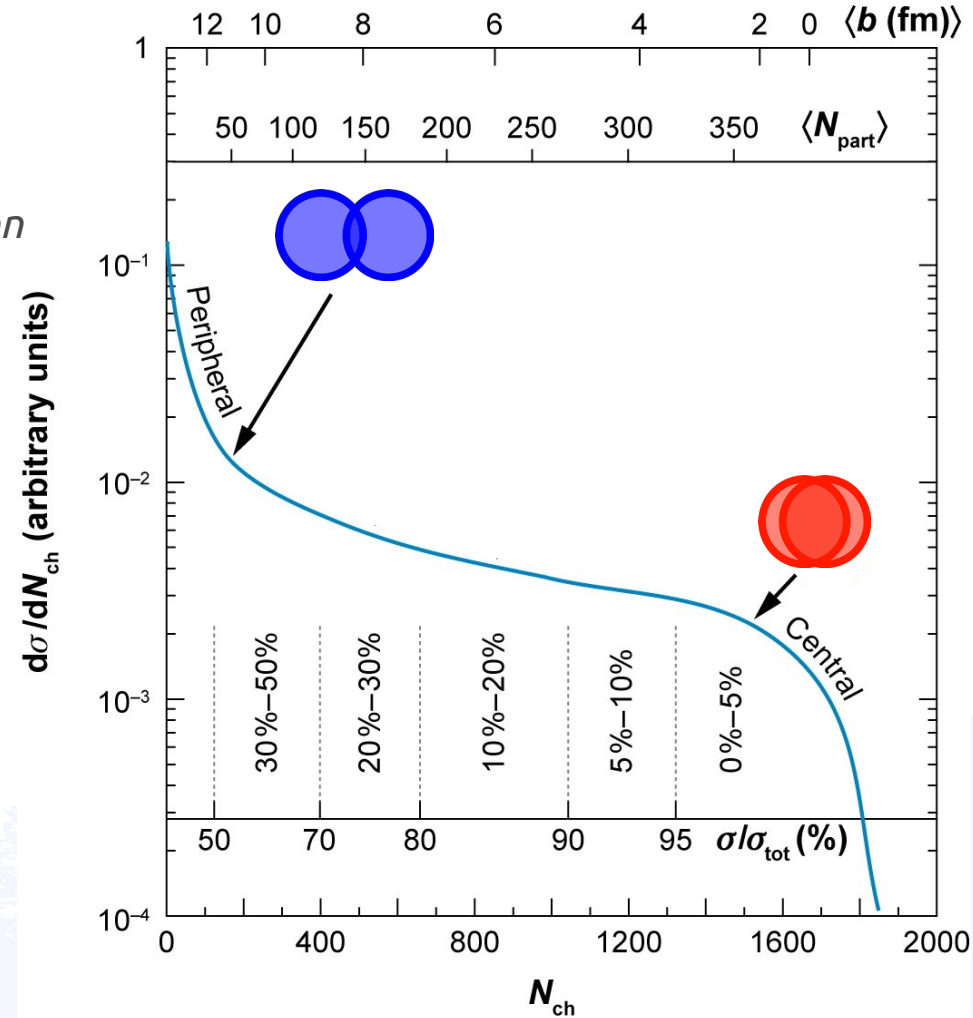


The thermodynamic properties of the medium can be derived from the measurement of final state particles (π , K, p, ...)

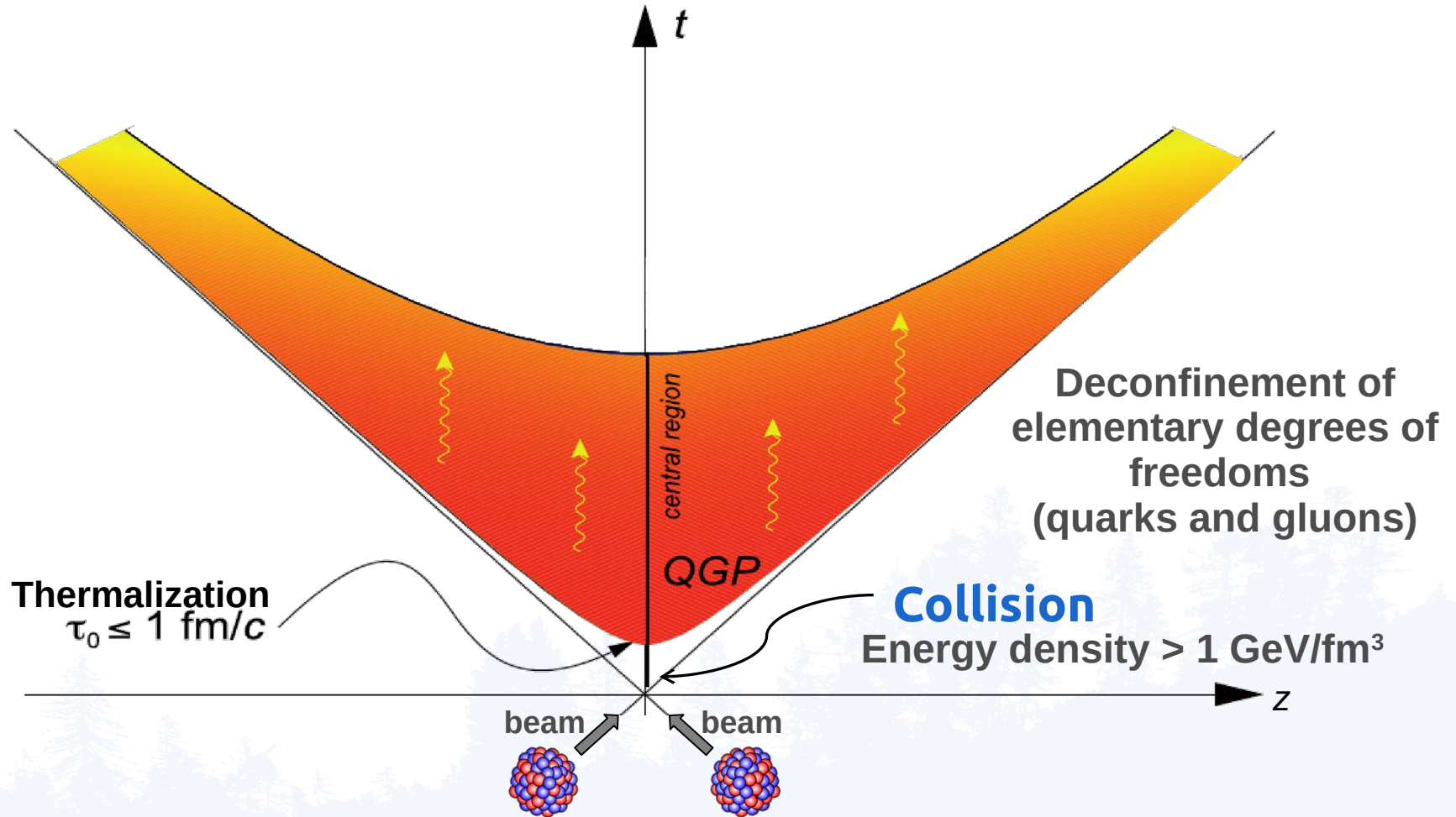
Heavy-ion collisions



- The impact parameter of the collision defines how many nucleons interact (N_{part}) and how many are not involved in the collision
- Centrality is expressed in fraction of the total inelastic cross section (σ_{tot})
- **The largest energy density is achieved in the most central collisions**



Evolution of a heavy ion collision

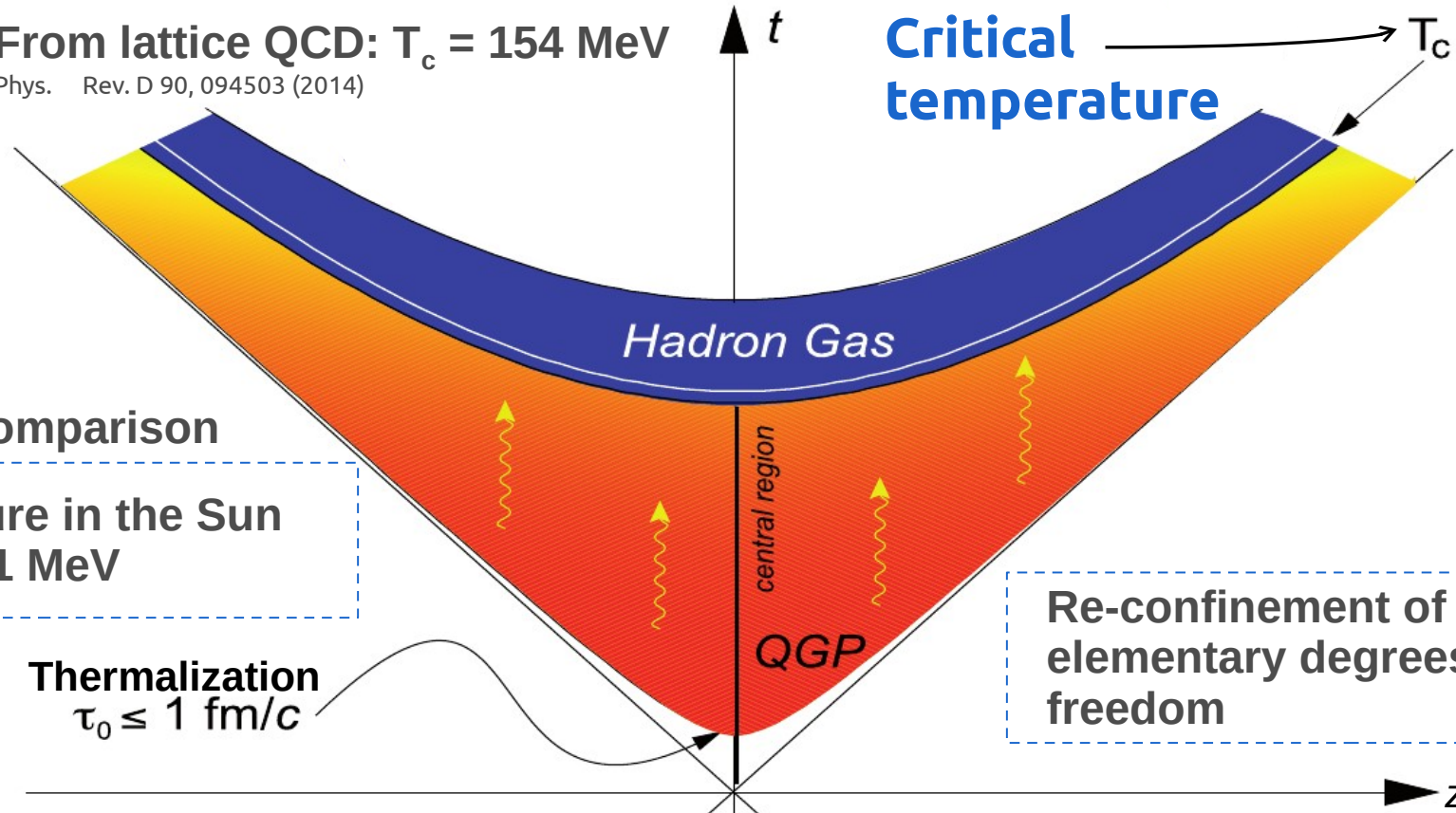


Evolution of a heavy ion collision

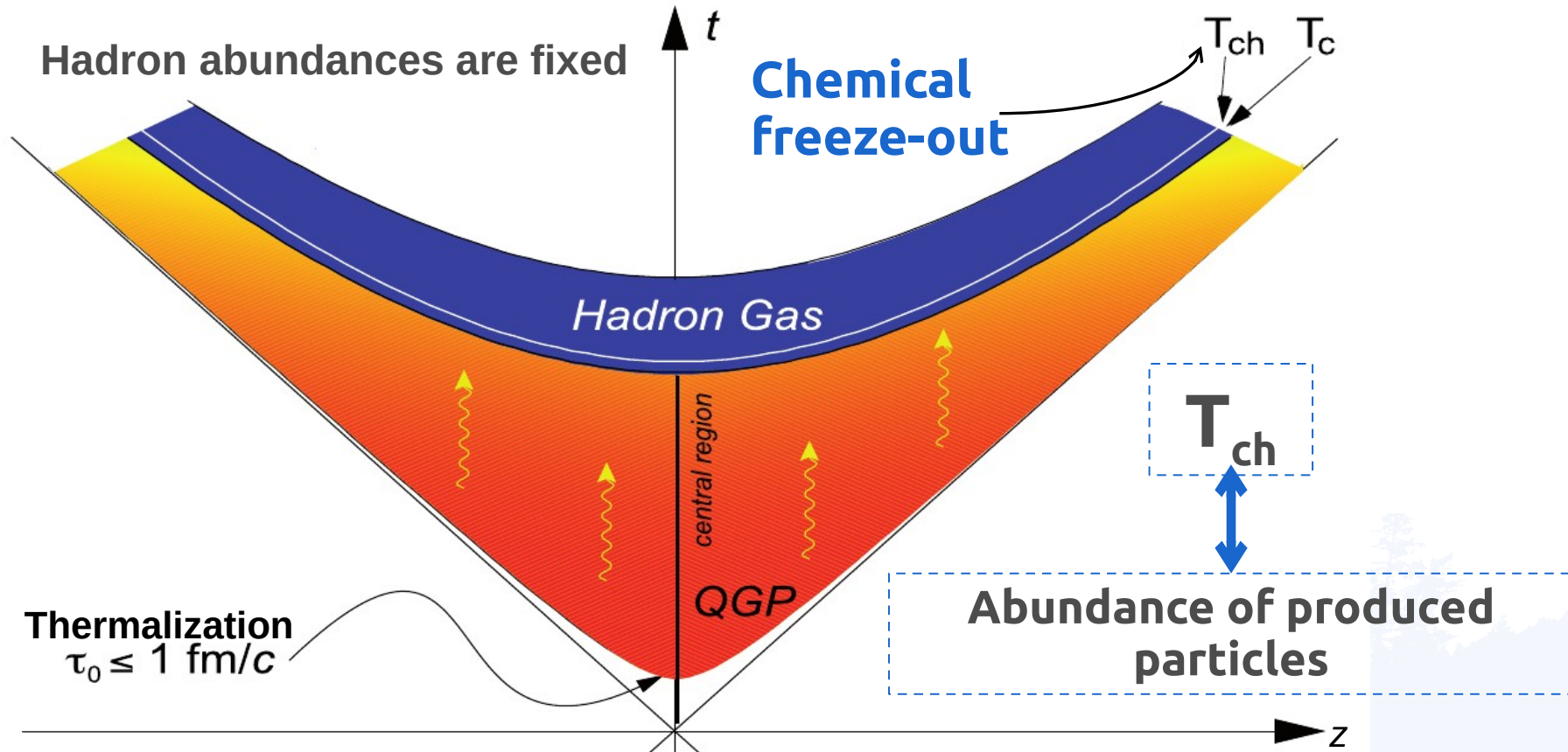
From lattice QCD: $T_c = 154 \text{ MeV}$

Phys. Rev. D 90, 094503 (2014)

Critical temperature T_c

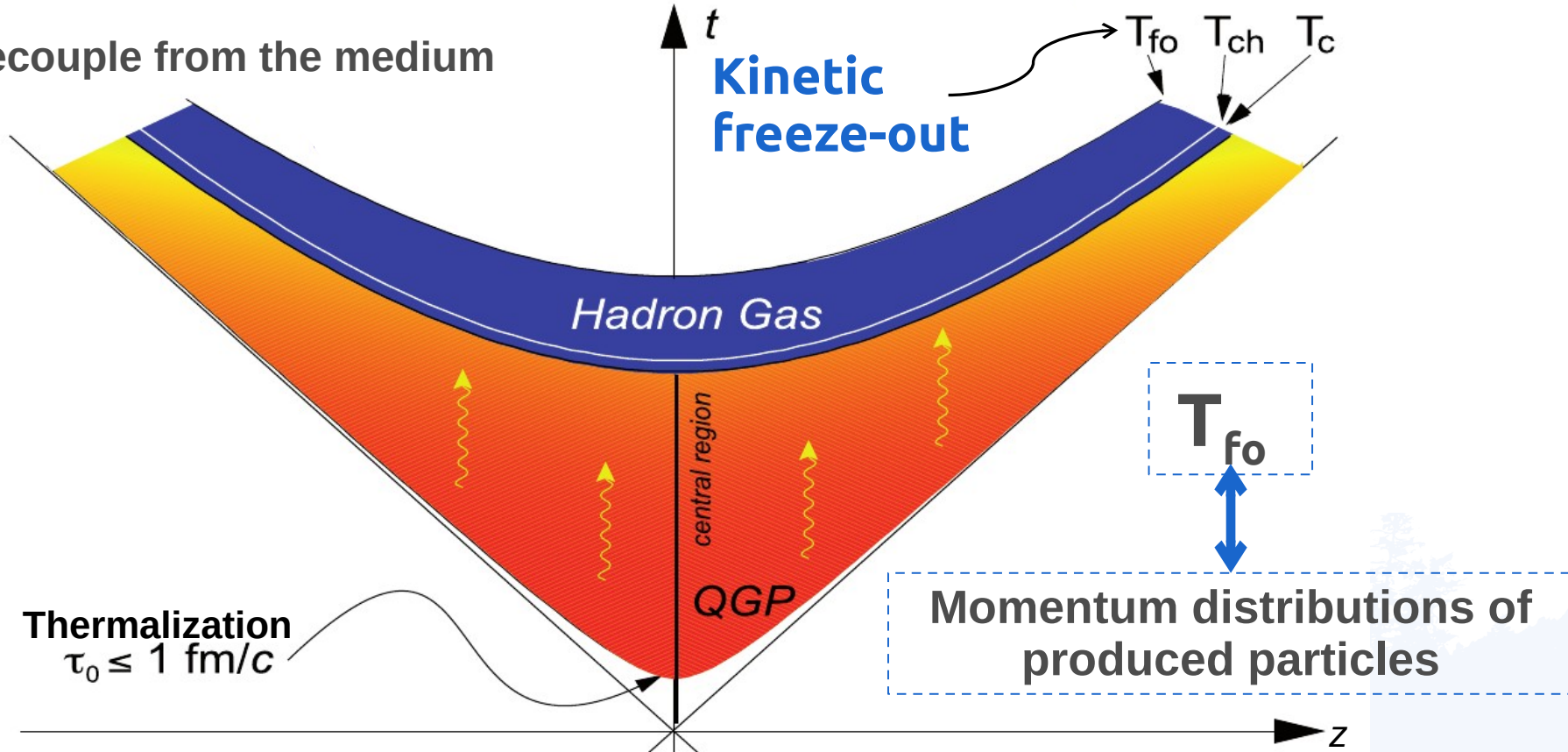


Evolution of a heavy ion collision



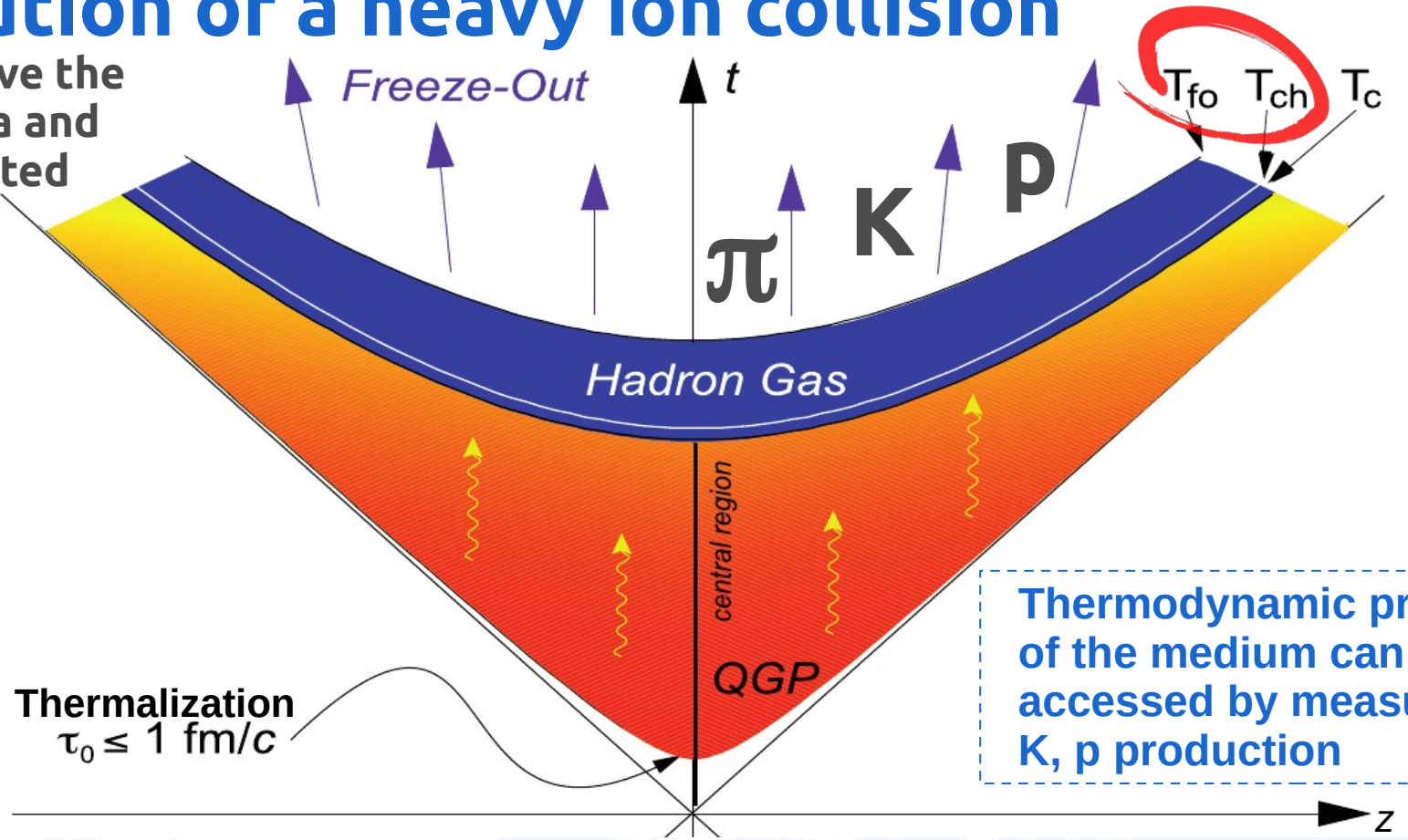
Evolution of a heavy ion collision

Hadrons decouple from the medium



Evolution of a heavy ion collision

Particles leave the collision area and can be detected

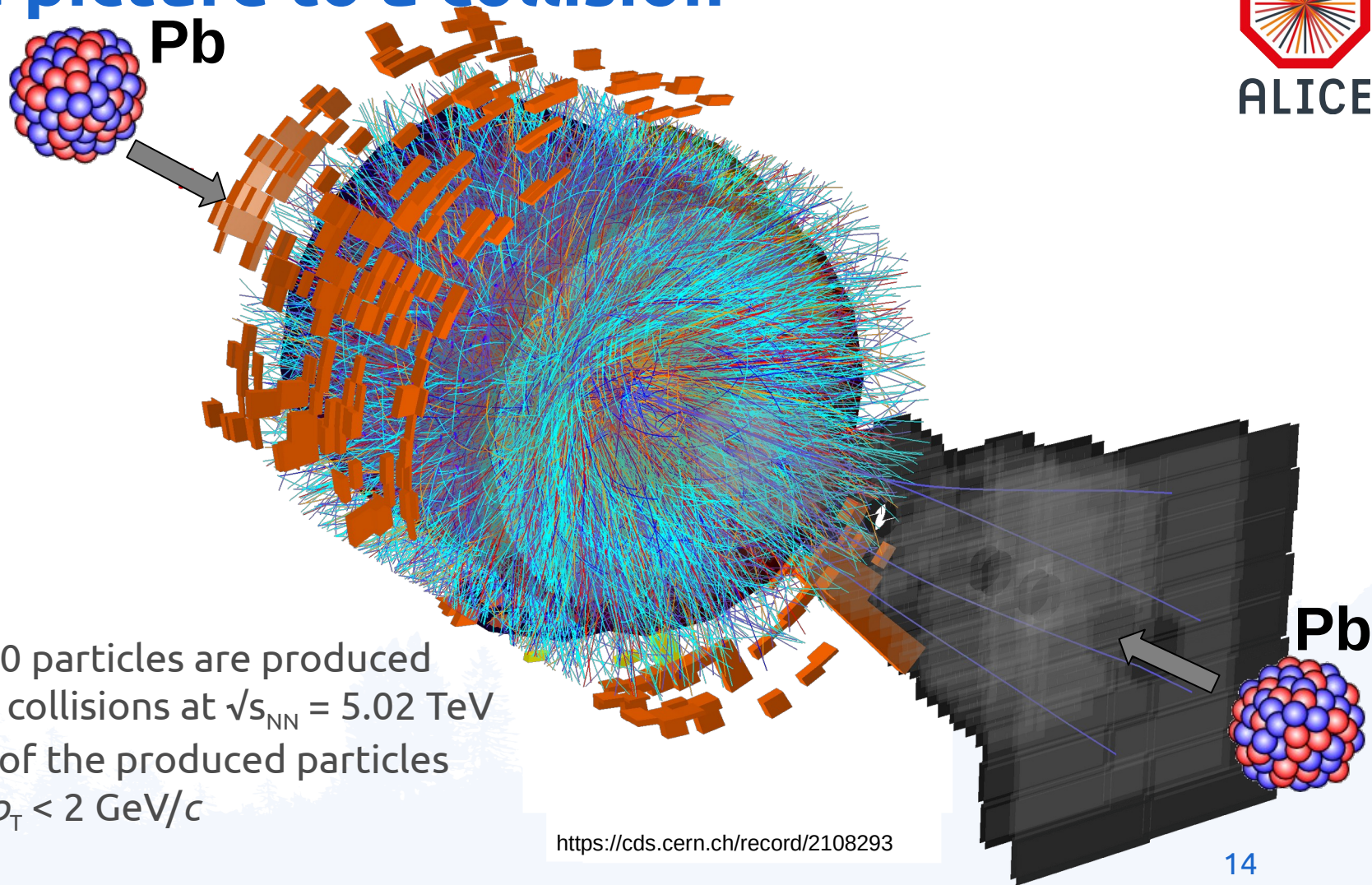


Thermodynamic properties of the medium can be accessed by measuring π , K , p production

Taking a picture to a collision



ALICE



- More than 20000 particles are produced in central Pb-Pb collisions at $\sqrt{s_{NN}} = 5.02$ TeV
- More than 98% of the produced particles are π , K, p with $p_T < 2$ GeV/c

<https://cds.cern.ch/record/2108293>

From large to small systems

Heavy ion collisions complemented by smaller collision systems

Nucleus-nucleus (AA)

- Deconfinement, QGP formation
- Testing QCD phase diagram
- Particle production ruled by thermodynamics and collectivity

Proton-nucleus (pA)

- Control experiment for AA
- Used to disentangle cold/hot nuclear matter effects
- Surprising features in high-multiplicity events

Proton-proton (pp)

- Baseline for both pA and AA collisions
- Similarities to AA in high-multiplicity event
- First pp collisions (900 GeV) of Run 3 at the end of last October

At the LHC:

Run 1 (2009-2013)	Run 2 (2015-2018)
pp 0.9, 2.76, 7, 8 TeV	pp 5.02, 13 TeV
p-Pb 5.02	p-Pb 5.02, 8.16 TeV
Pb-Pb 2.76 TeV	Pb-Pb 5.02 TeV Xe-Xe 5.44 TeV

Whole-lotta particles to measure

Light flavor hadrons
have only u, d, s quarks

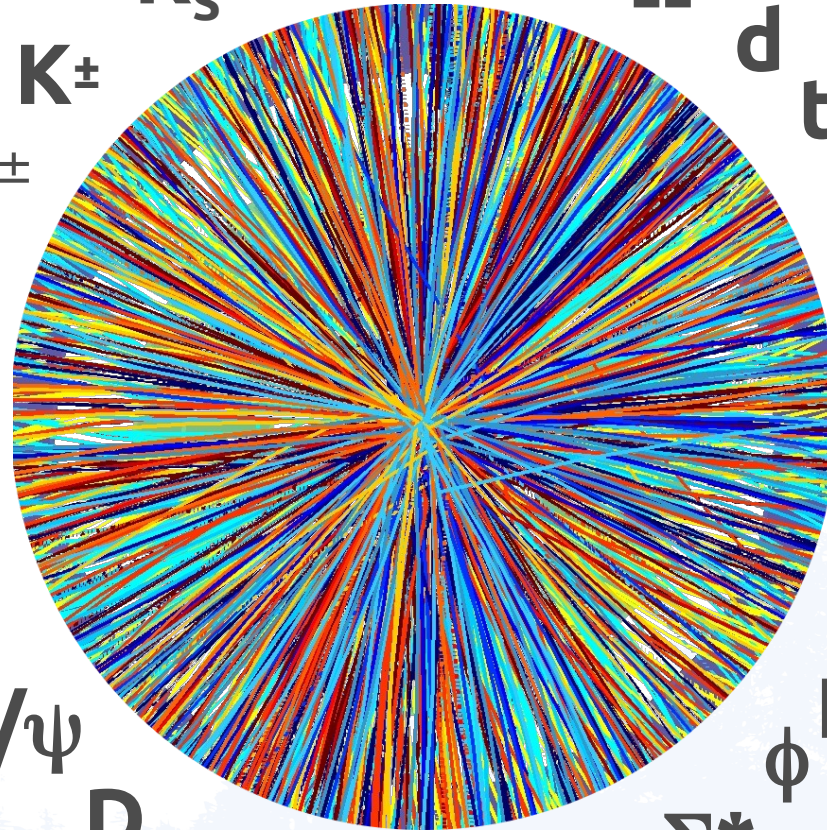
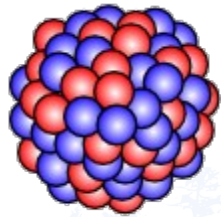
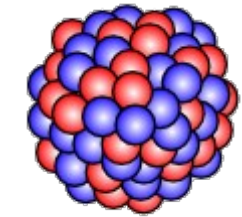
Strange content

Light nuclei

Hyper-nuclei

Resonances

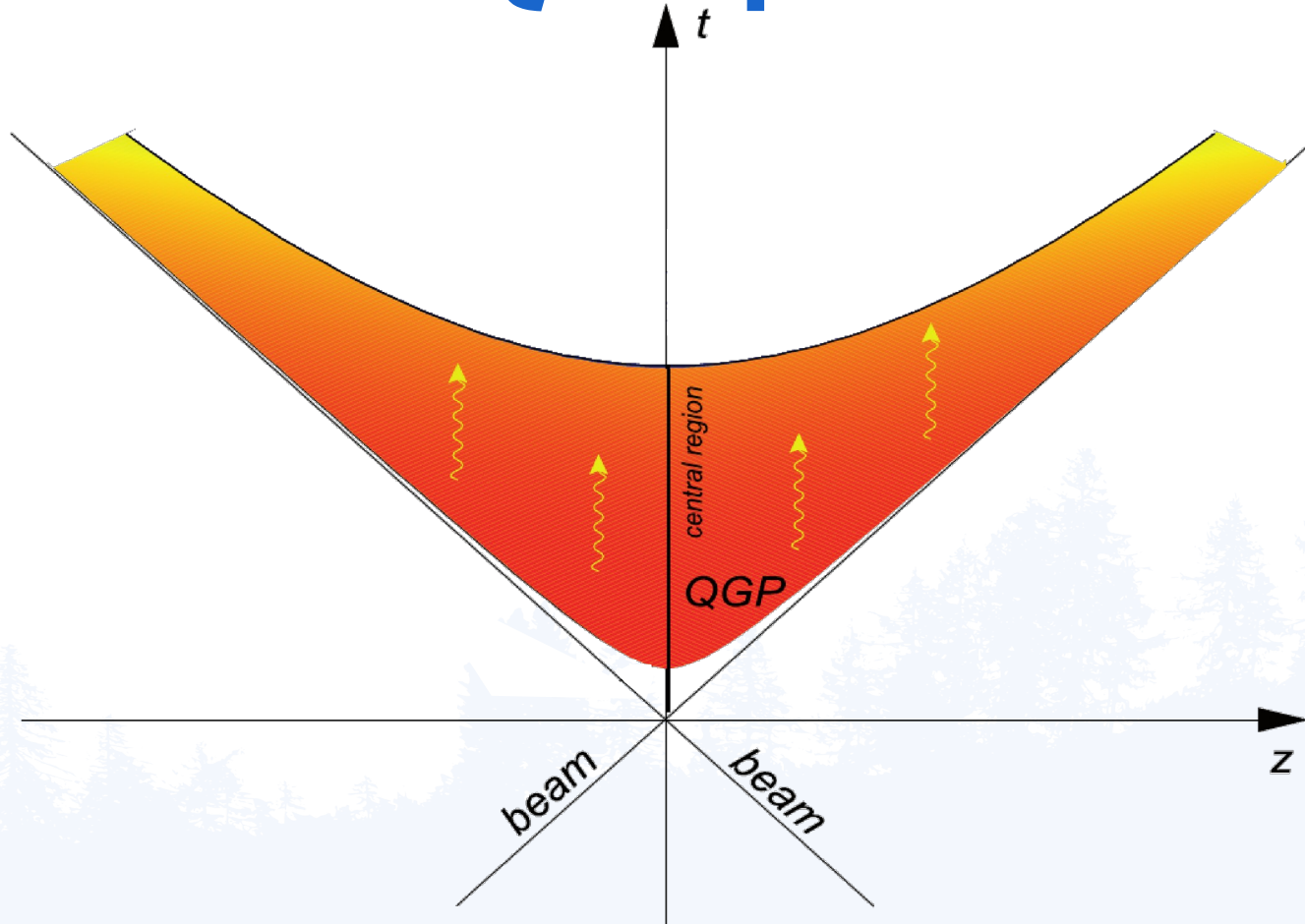
Heavy flavor
hadrons have
also *heavy* quarks



Not exhaustive!

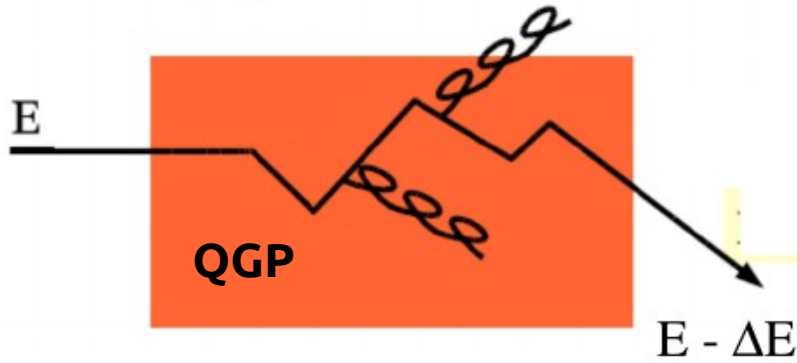
K_s^0 p Λ Ξ Ω d t
 $K^\pm$ $\pi^\pm$ π^0 Λ_c Y ψ' J/ψ D_s D^{*+} D^0 Ξ^* Λ^* Σ^* ϕ K^*

Investigating the QGP phase



Nuclear modification factor: R_{AA}

- Useful information on the QGP can be obtained from hard probes i.e. highly energetic partons produced in hard scatterings



Partonic energy loss in the medium due to **gluon radiation** and **collisions** with the medium partons

$$R_{AA} = \frac{(d^2N/dydp_T)_{AA}}{\langle N_{coll} \rangle (d^2N/dydp_T)_{pp}}$$

- No QGP is expected to form in pp collisions
- R_{AA} quantifies the difference between Pb-Pb collisions and the sum of $\langle N_{coll} \rangle$ incoherent collisions i.e. quantifies the effect of the presence of the medium

Nuclear modification factor: R_{AA}

R_{AA} accounts for the modification of the medium on the particle spectrum

– $R_{AA} < 1$:

particles are absorbed or lose their energy in a medium opaque to the color charge (QGP)

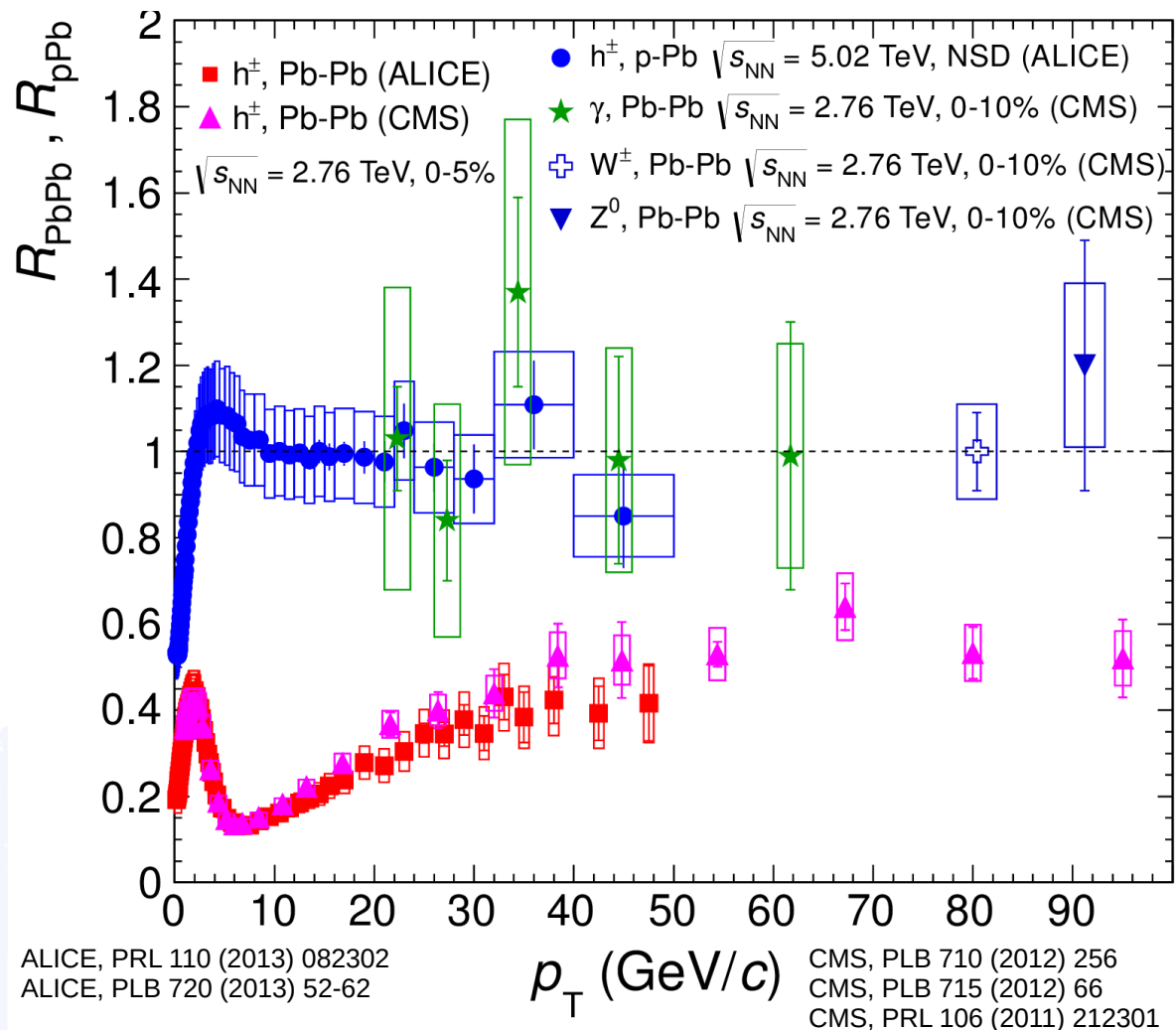
– $R_{AA} = 1$:

the presence of a dense medium cannot be seen on the produced particles

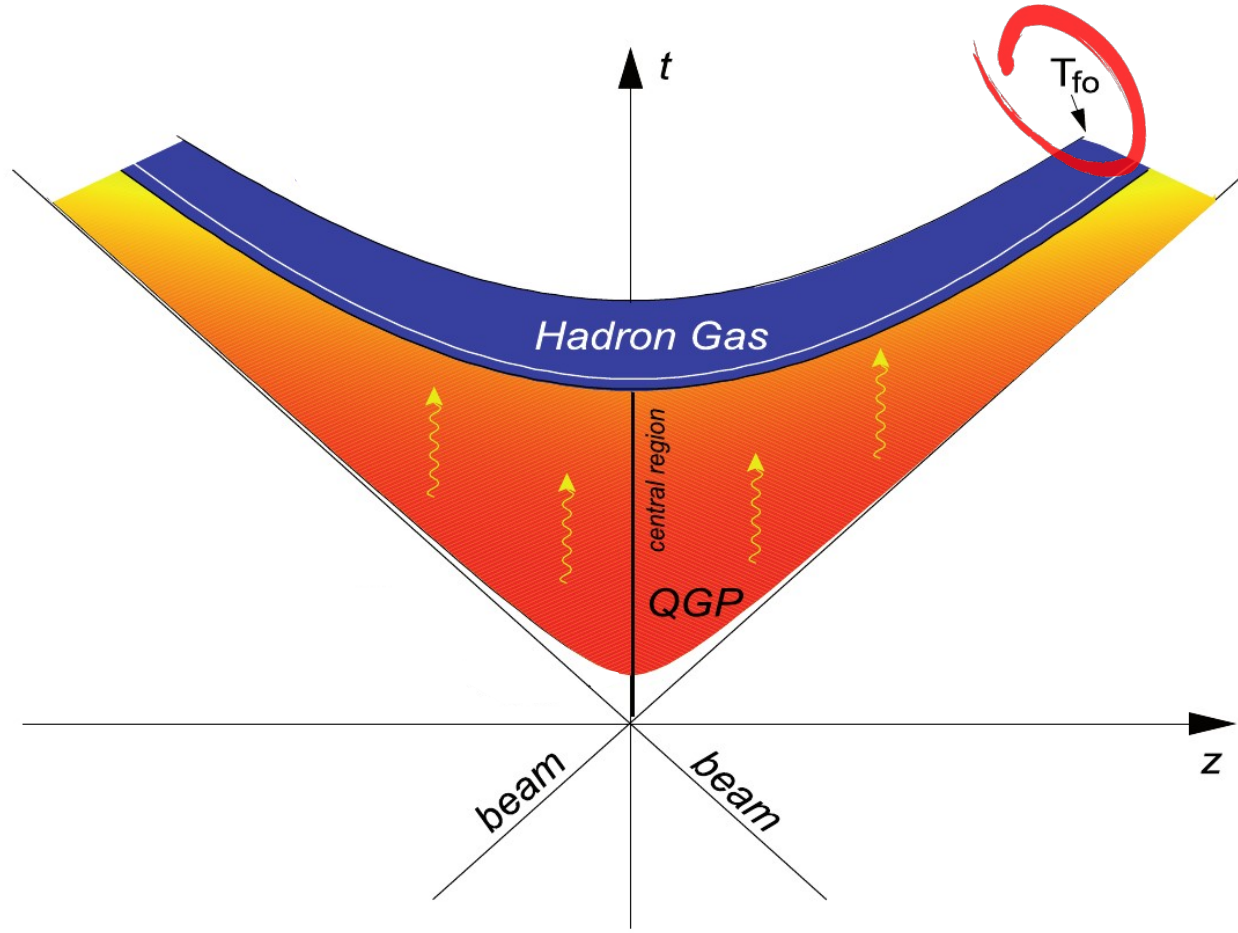
– Only color charged probes are affected by the presence of the medium

– γ , W and Z bosons are unaffected by the medium as they cannot lose energy via gluon radiation

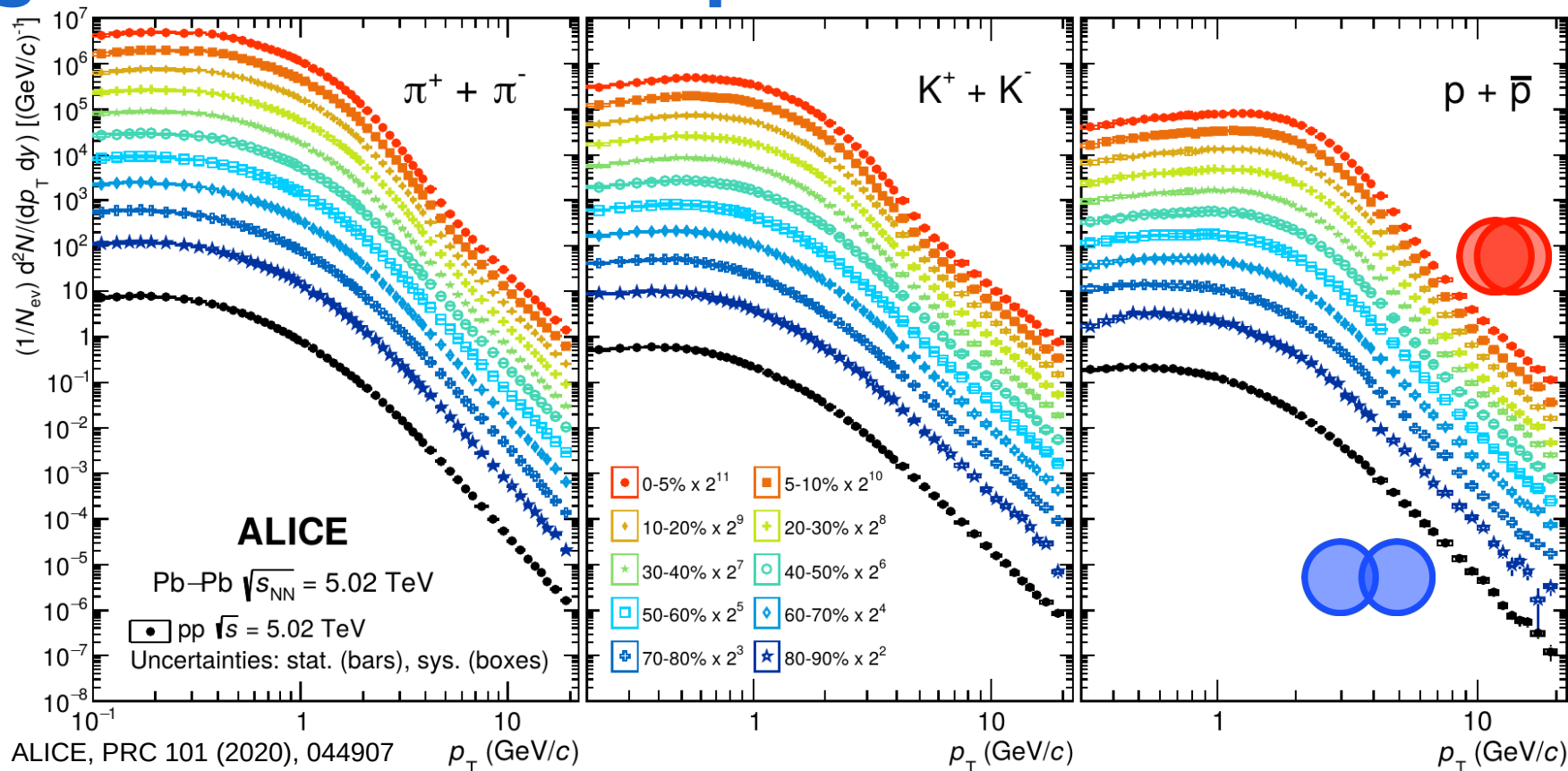
Different hadrons gives information of quark interaction with medium



Measuring the system at the kinetic freeze-out



Light flavor hadron spectra



- **Mass-dependent hardening** of the soft part with increasing centrality due to the **collective expansion (*radial flow*)**
- **Depletion at low p_T and enhancement at intermediate p_T**

Blast-Wave model

The Blast-Wave model describes the particle distribution at the kinetic freeze-out as a result of the expansion of a thermalized source

- The Boltzmann-Gibbs statistics is used as an initial thermal distribution
- The expanding source causes a mass dependent hardening
- The expansion velocity and decoupling temperature are free parameters of the model

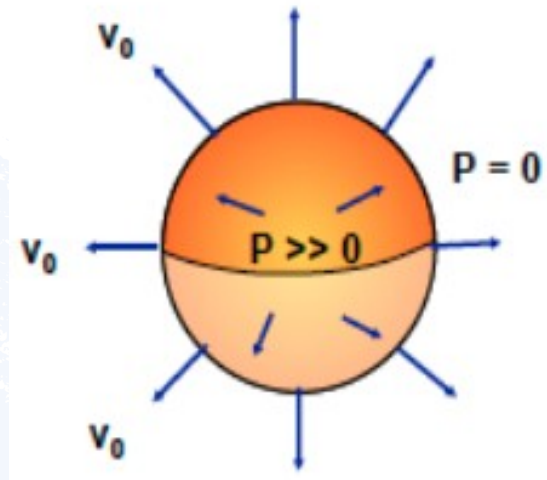
$$E \frac{d^3 N}{d p^3} \propto \int_0^R m_T I_0 \left(\frac{p_T \sinh(\rho)}{T_{Kin}} \right) K_1 \left(\frac{m_T \cosh(\rho)}{T_{Kin}} \right) r dr$$

$$m_T = \sqrt{m^2 + p_T^2} \quad \rho = \tanh^{-1}(\beta_T) \quad \beta_T = \beta_s \left(\frac{r}{R} \right)^n$$

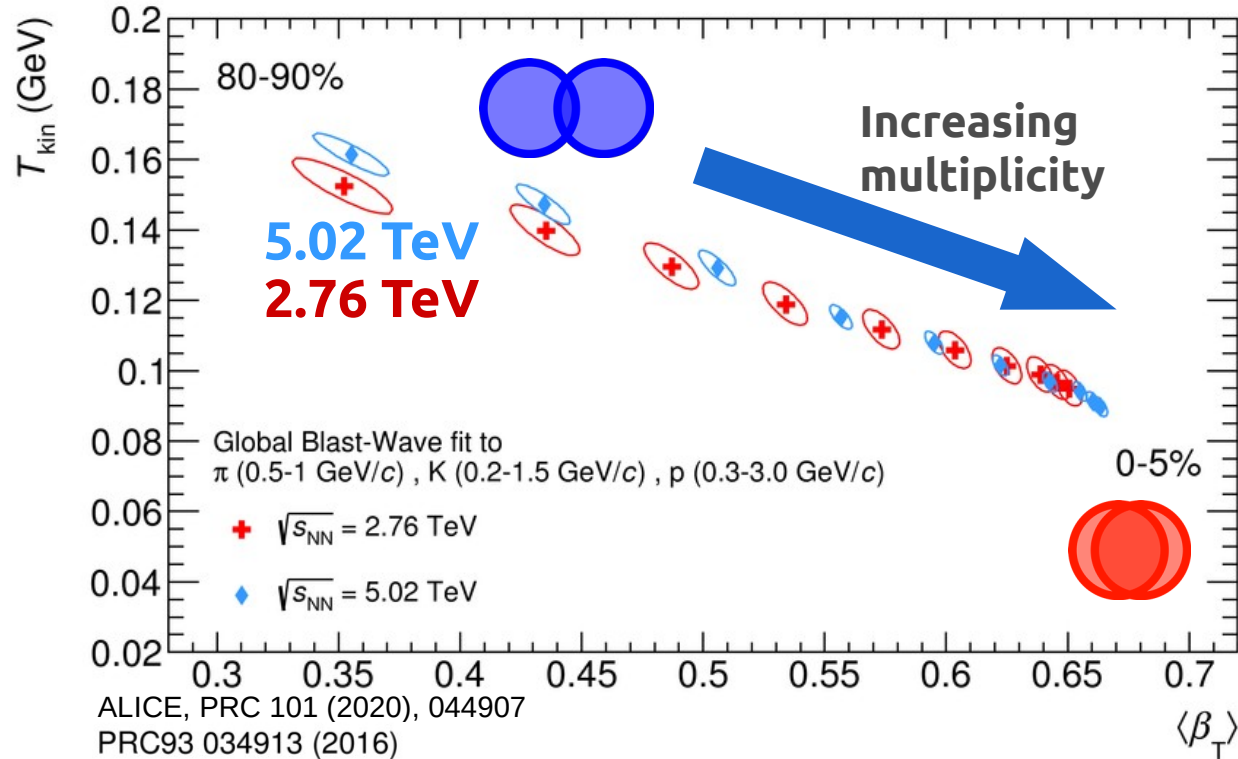
Schnedermann, Sollfrank and Heinz Phys. Rev. C 48, 2462

» $\beta_T \rightarrow$ radial expansion velocity

» $T_{Kin} \rightarrow$ kinetic freeze-out



Blast-Wave model: β_T vs T_{kin}

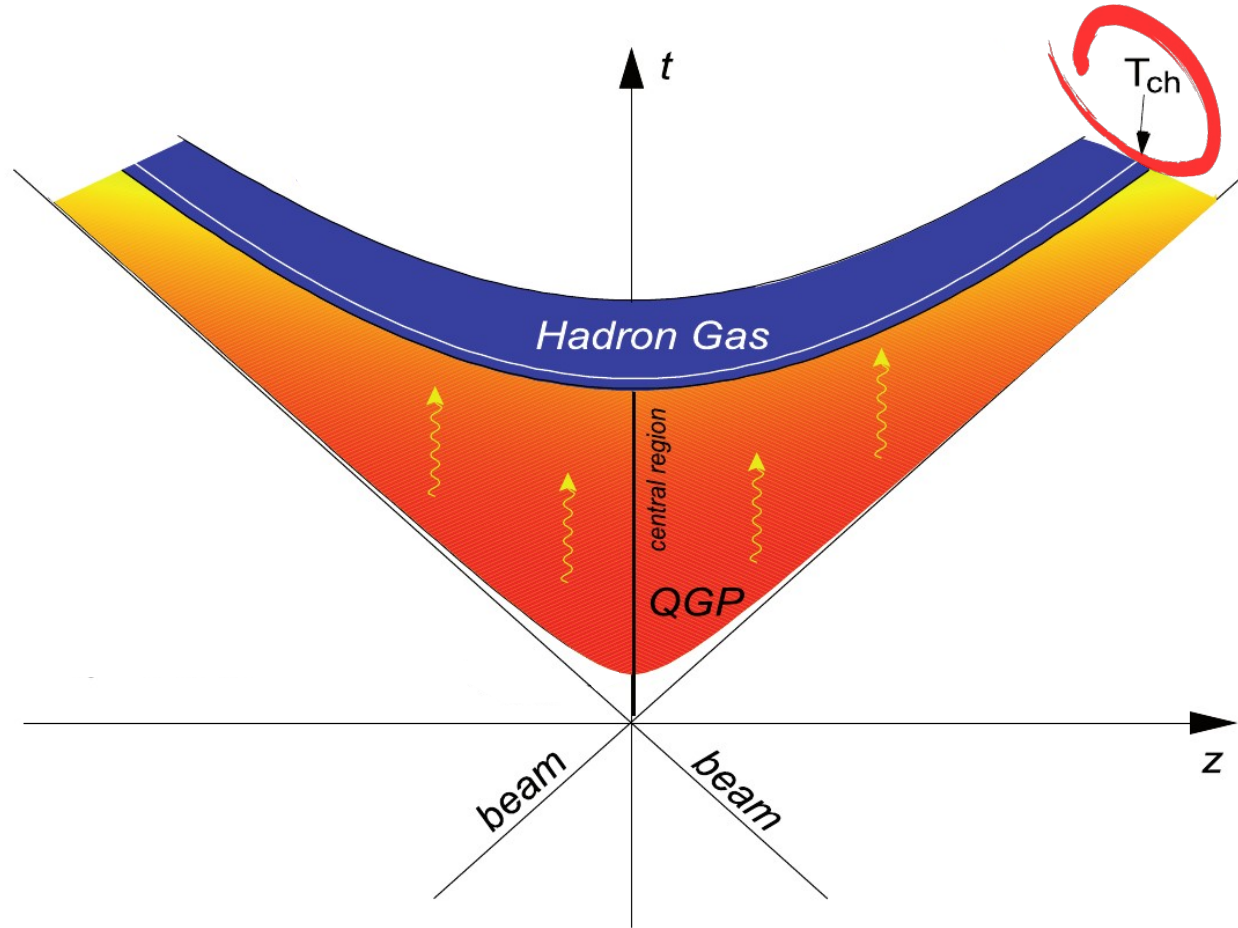


- The free parameters of the BW model are obtained with a simultaneous fit to the π , K, p p_T -distributions

- Maximum expansion velocity is reached in the most central collisions
- The expansion velocity slightly increases with the collision energy

- Central collisions exhibit the lowest kinetic freeze-out temperature (~ 85 MeV)
- The temperature decreases with increasing collision energy $\rightarrow$ longer lived system?

Measuring the system at the chemical freeze-out



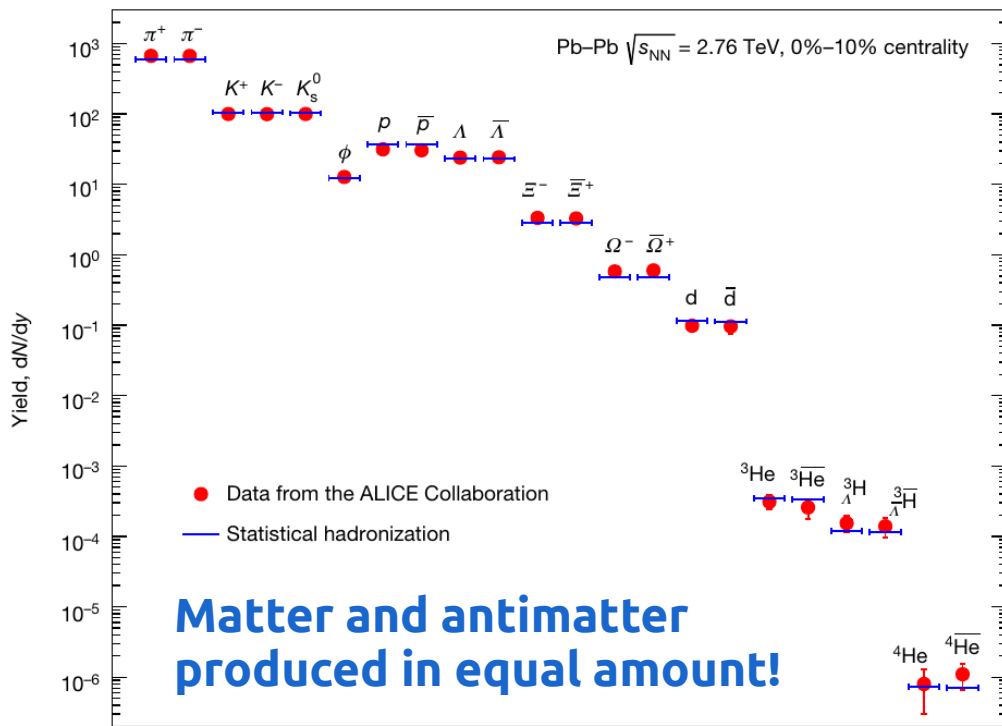
Thermal model

- At the chemical freeze-out, the system (hadron resonance gas) is in thermal and chemical equilibrium
- The particle abundances in a thermalized medium can be derived as a function of its thermodynamic properties (temperature and volume) by writing the system's partition function
- In heavy-ion collisions the grand-canonical ensemble is used

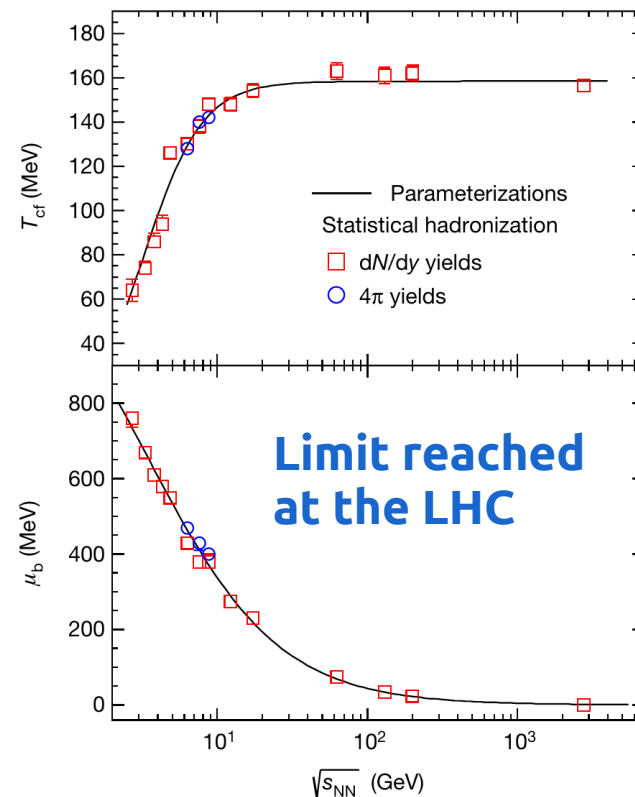
$$\ln Z^{GC}(T, V, \{\mu_i\}) = \sum_{\text{species } i} \frac{g_i V}{(2\pi)^3} \int d^3p \ln \left(1 \pm e^{-\beta(E_i - \mu_i)} \right)^{\pm 1} \rightarrow N_i^{GC} = T \frac{\partial \ln Z^{GC}}{\partial \mu_i}$$

- The quantum number conservation (baryon number, strangeness, electric charge) in the reaction is ensured by chemical potential μ_i , that can be fixed from the quantum number of the initial stage

Thermal model to describe particle yields

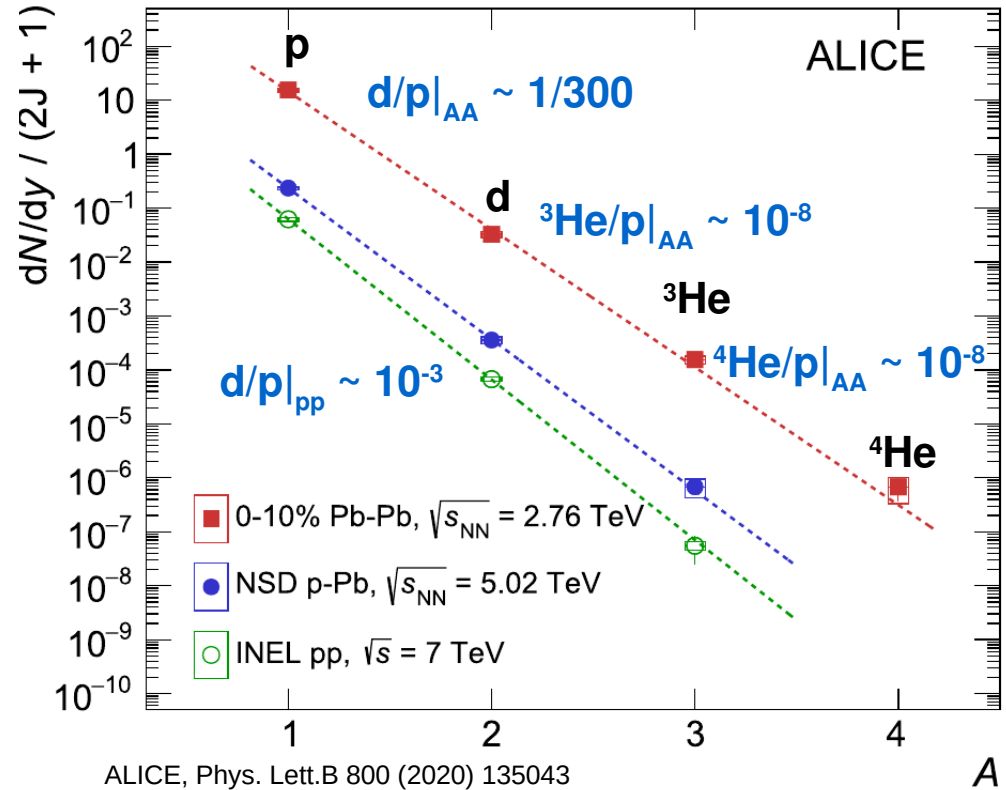
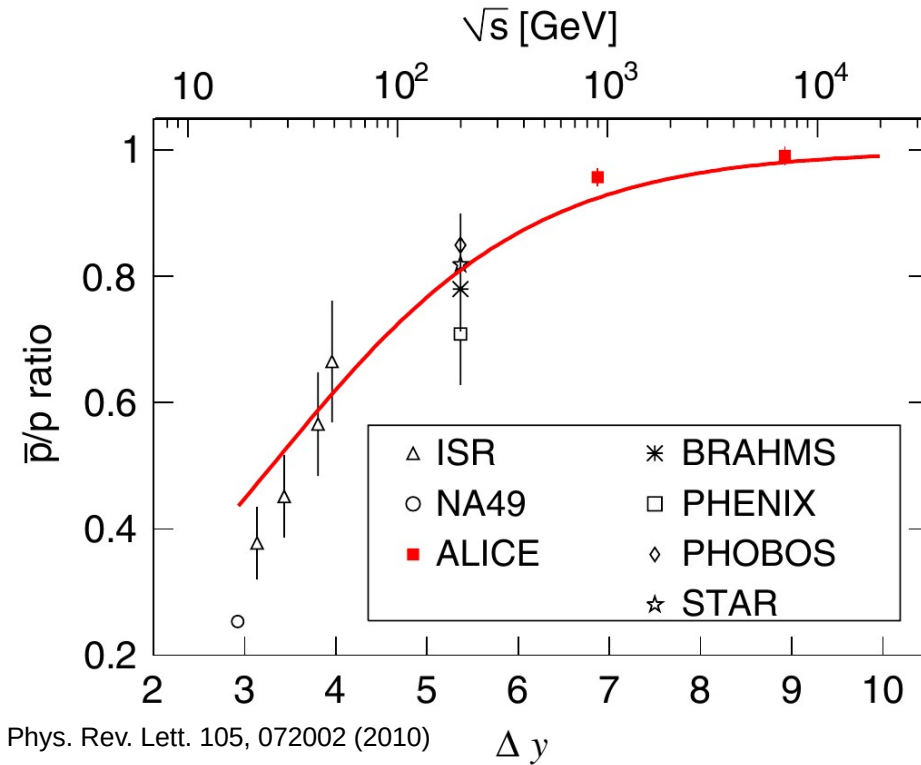


Andronic et al., Nature volume 561, pages 321–330 (2018)



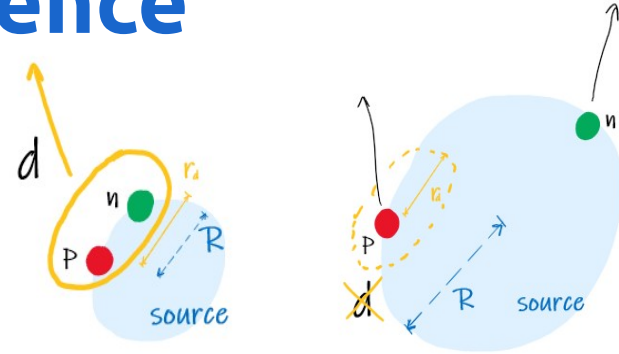
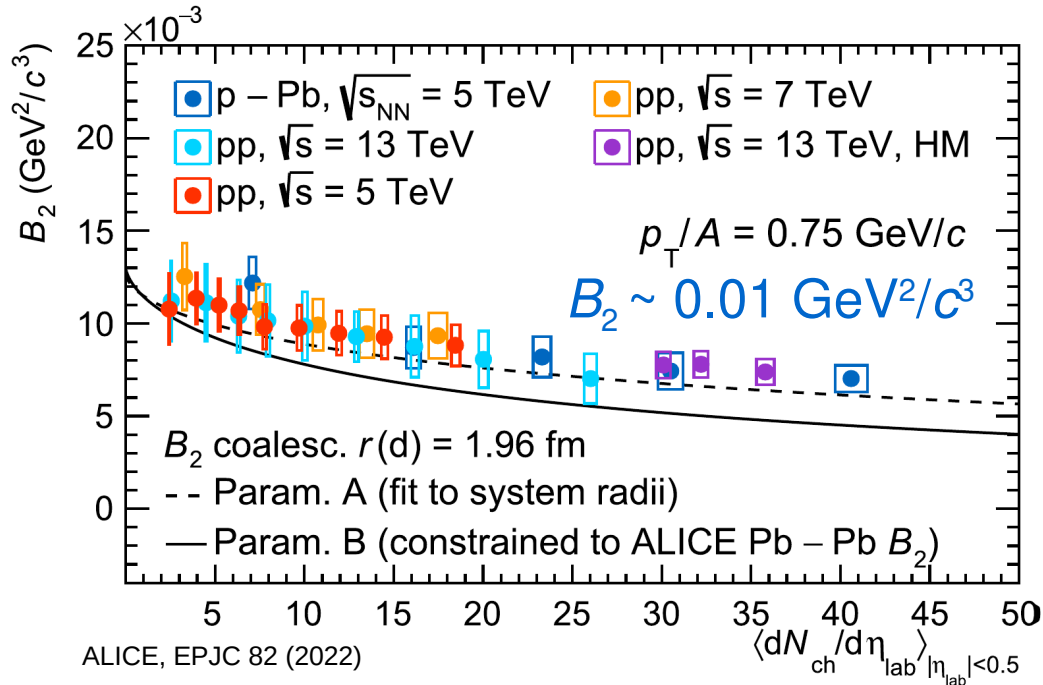
- Single chemical freeze-out temperature for all particle species (common source) in central Pb-Pb collisions $T_{ch} \sim 156 \pm 1.5$ MeV
- This value is in close to the critical temperature T_c (~ 154 MeV) obtained from lattice QCD
→ **phase transition is close to chemical freeze-out**

Antimatter factories with heavy ions



- Matter-antimatter ratio at LHC ~ 1 **independently on the collision system**
- Production rates strongly depend on the mass number and collision system

Nuclei formation by coalescence



$$E_A \frac{d^3 N_A}{dp_A^3} = B_A \left(E_{p,n} \frac{d^3 N_{p,n}}{dp_{p,n}^3} \right)^A \bigg|_{\vec{p}_p = \vec{p}_n = \frac{\vec{p}_A}{A}}$$

For a nucleus large w.r.t. the source
 → “simple” coalescence models

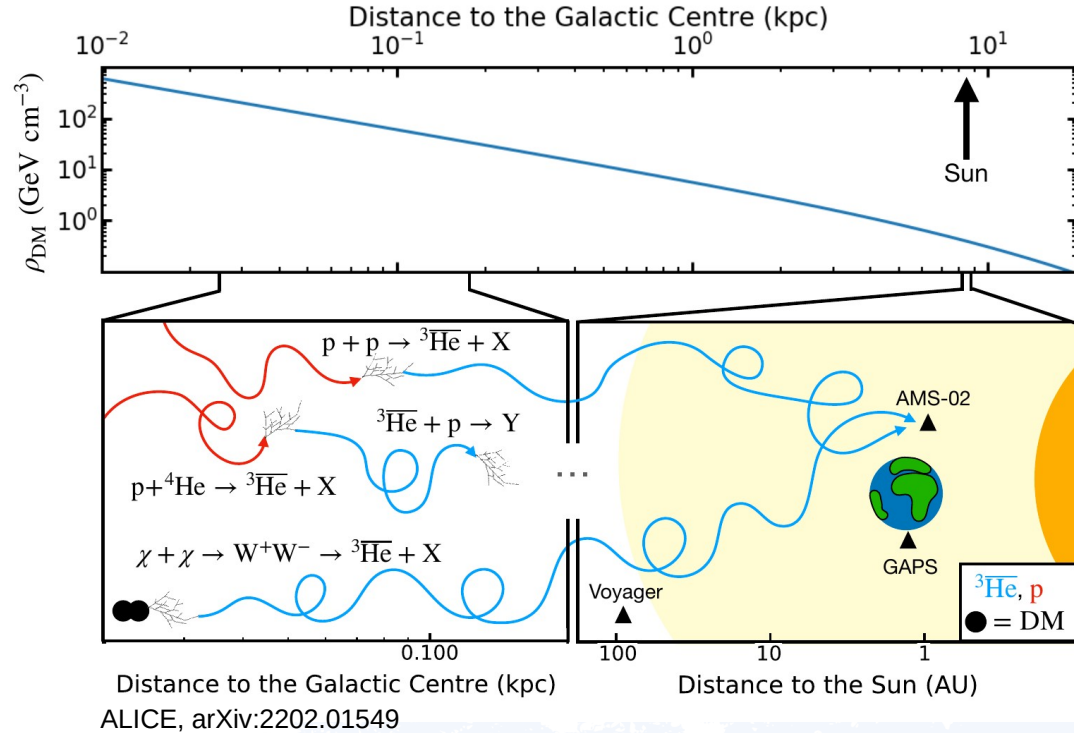
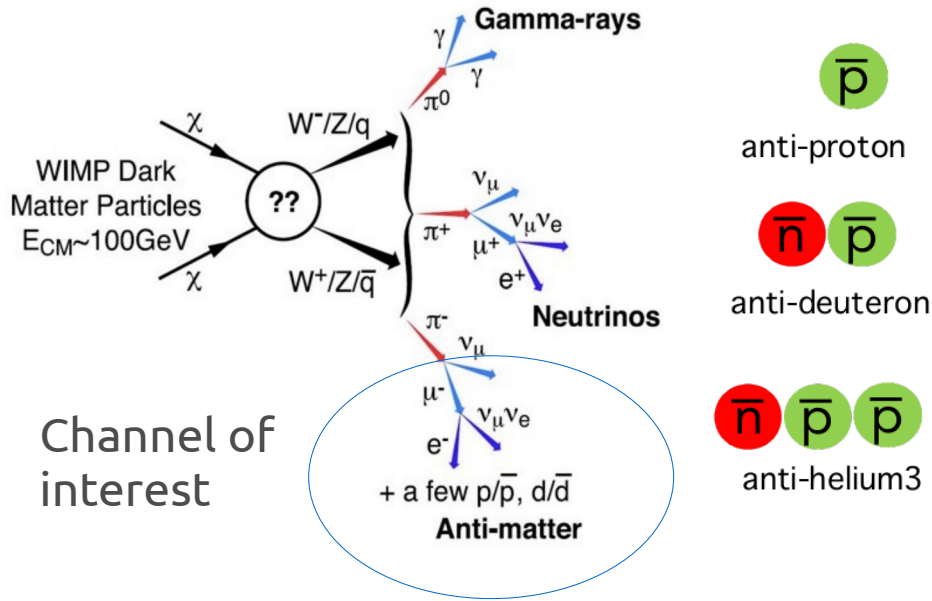
$$B_A = \left(\frac{4\pi}{3} p_0^3 \right)^{(A-1)} \frac{1}{A!} \frac{M}{m^A}$$

$\leftarrow$ Nucleus mass
 $\leftarrow$ Nucleon mass

Coalescence momentum

- (Anti-)nuclei can be formed at kinetic freeze-out by coalescence of nucleons close enough in phase space
- Production depends on the **coalescence probability B_A**
- **B_A measured extensively in small and large systems**

Antimatter as dark matter “smoking gun”



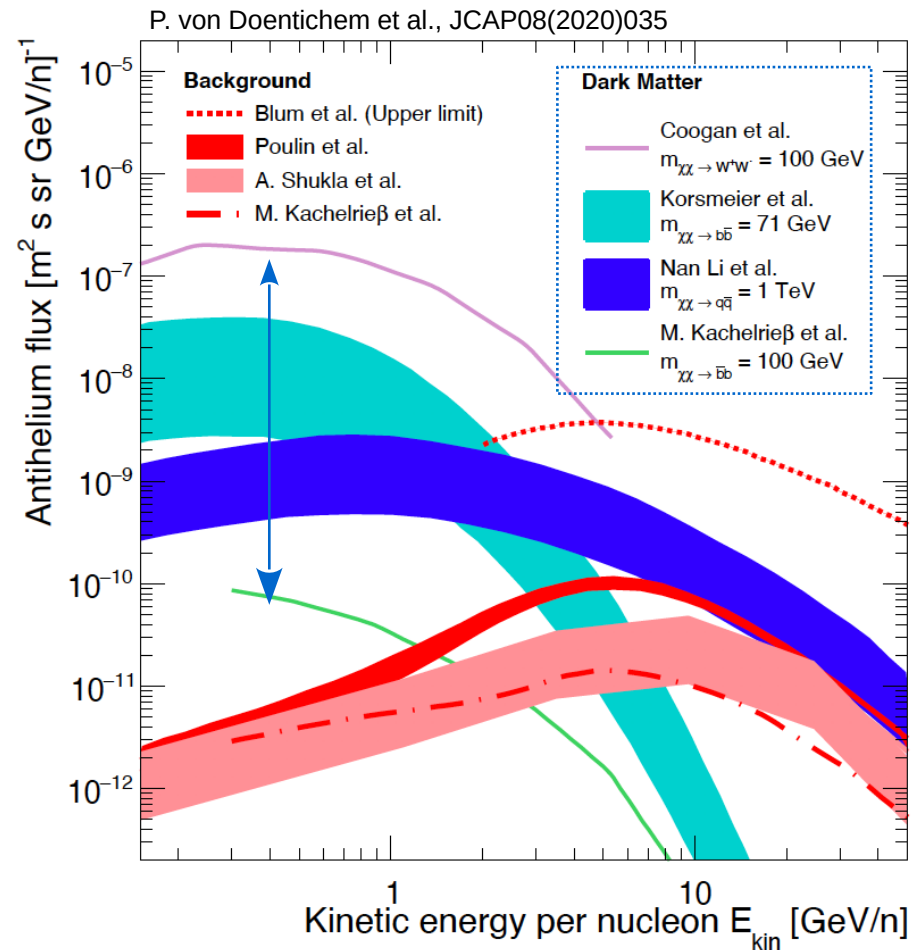
- **Anti-p and anti-n are produced** by WIMP annihilation into SM channels
- Anti-deuterons and anti- ${}^3\text{He}$ are produced via coalescence of anti-nucleons
- **No anti-nuclei as primary cosmic rays (only nuclei $\rightarrow$ mostly protons and helium)**
- Secondary anti-p, anti-d, anti- ${}^3\text{He}$ produced by interaction of primary CR with the InterStellar Matter (pp , $p\text{-He}$, ...) $\rightarrow$ **background for the DM signal**

The dark matter source

Anti-nuclei flux from Dark Matter **depends on the details of the particle physics model** and the DM density in a given point of the Galaxy

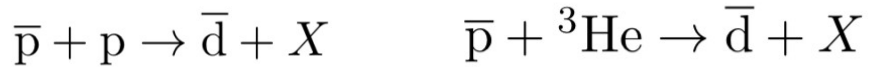
In particular:

- thermally-averaged annihilation cross section into SM channel
- DM mass, e.g. $70 < m_{\text{DM}} < 100 \text{ GeV}/c^2$
- energy spectrum of the products
- DM density in the vicinity of the solar system



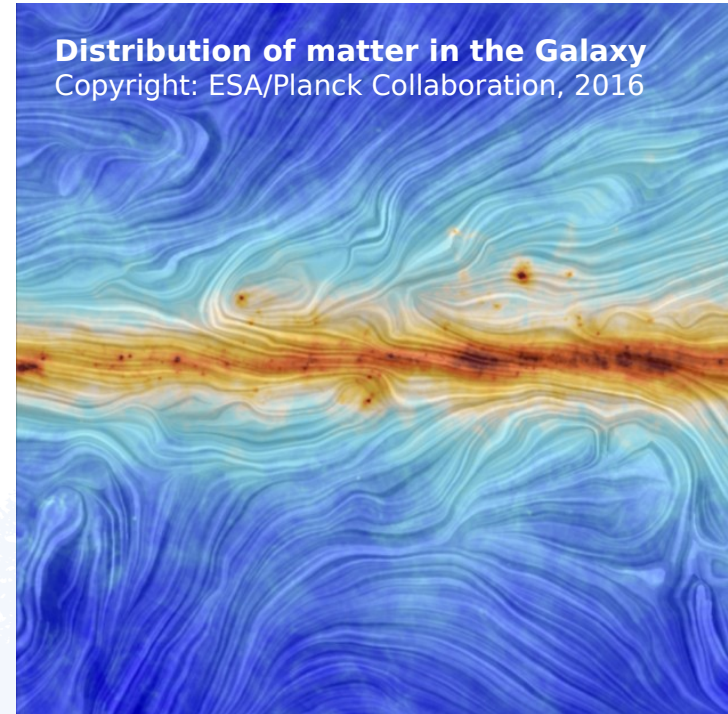
The background source

Secondary anti-nuclei can be produced by spallation reactions of primary CR with ISM, e.g.



- **Depends on the cross-sections for $\bar{p}$ production in pp, p-He, p-A**
- Achieved via the coalescence mechanism (as for DM) $\rightarrow$ different anti-nucleon distributions
coalescence momentum unknown

In addition, a tertiary CR component:

$$\bar{d} + p \rightarrow \bar{d} + X$$
$$\bar{d} + {}^3\text{He} \rightarrow \bar{d} + X$$


Towards precise estimate of anti-nuclei flux

Anti-matter flux predictions for signal from DM and background from secondary/tertiary CR require as input:

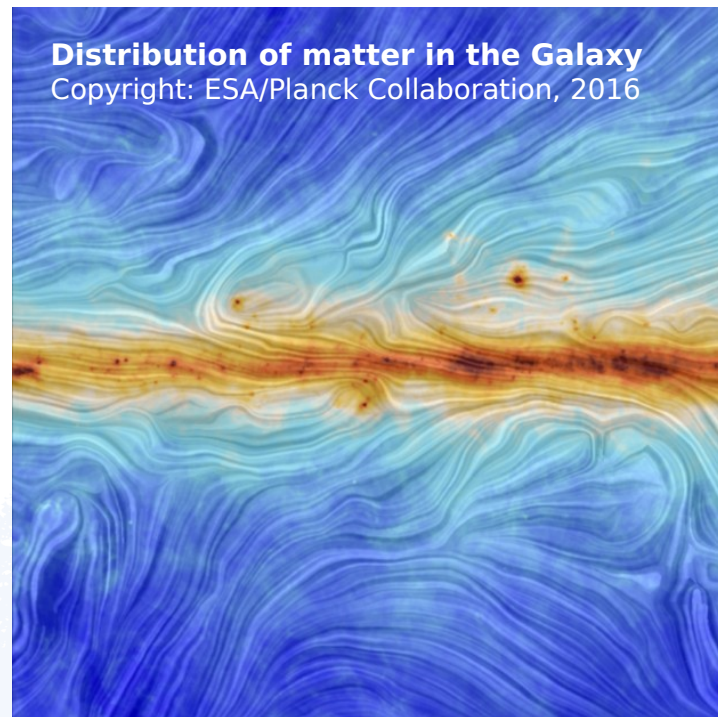
- **Coalescence mechanism**
- **Coalescence momentum**
- **Cross-sections for anti-p production in pp, p-He, p-A**

To measure at the collider!

Flux calculations sensitive to the astrophysical details, i.e. how particles propagate in the Galaxy:

- Acceleration by Super Novae remnants
- Diffusion in the galactic magnetic field ($\sim \mu\text{Gauss}$)
- Energy loss / gain (*for loosely bound nuclei, break-up dominates*)
- Solar modulations (*matter mostly at low E , where DM signal prominent*)

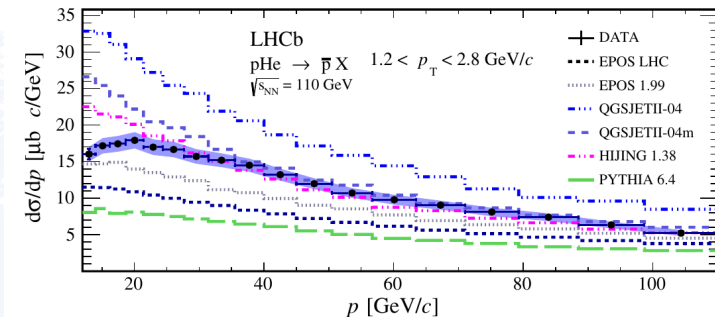
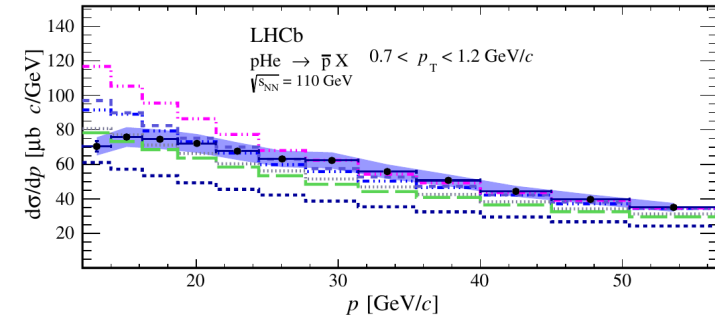
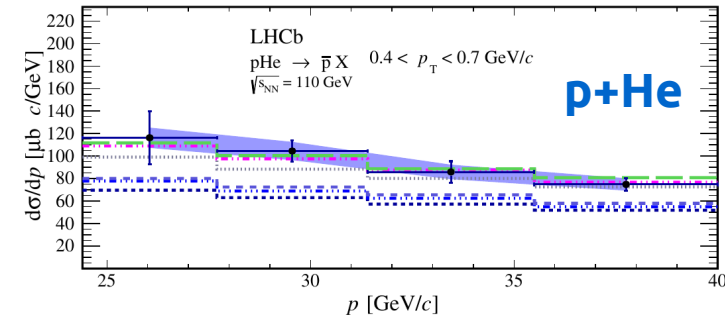
Tuned on astrophysical measurements



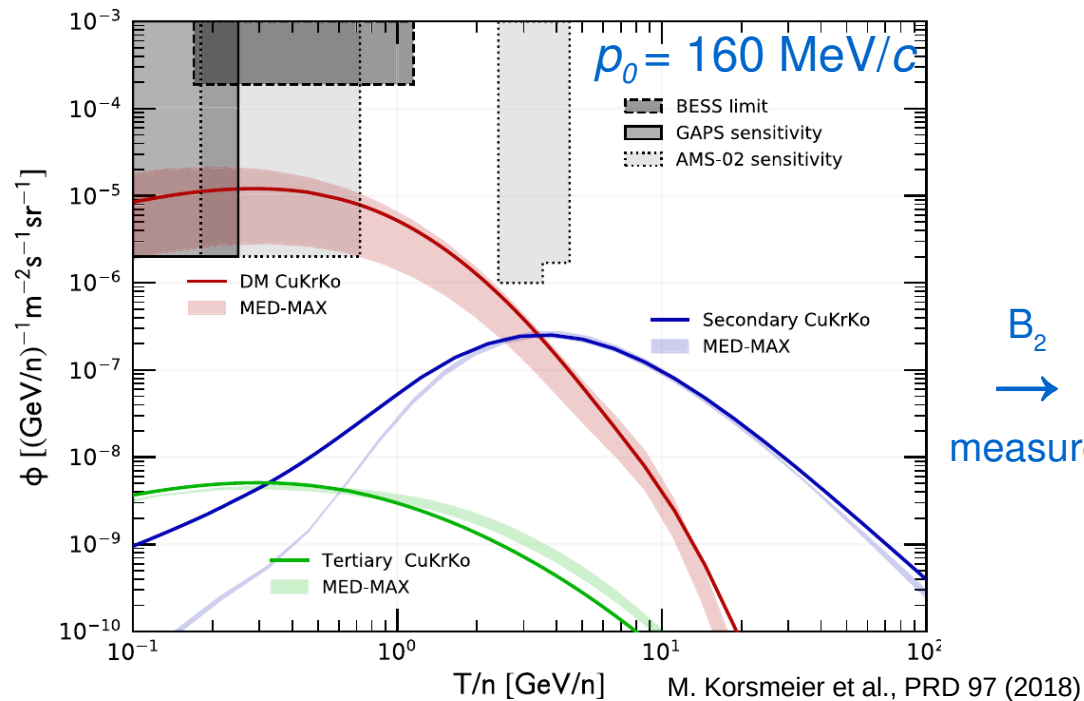
Constraining background to DM searches

Fixed target experiment with the LHCb SMOG system

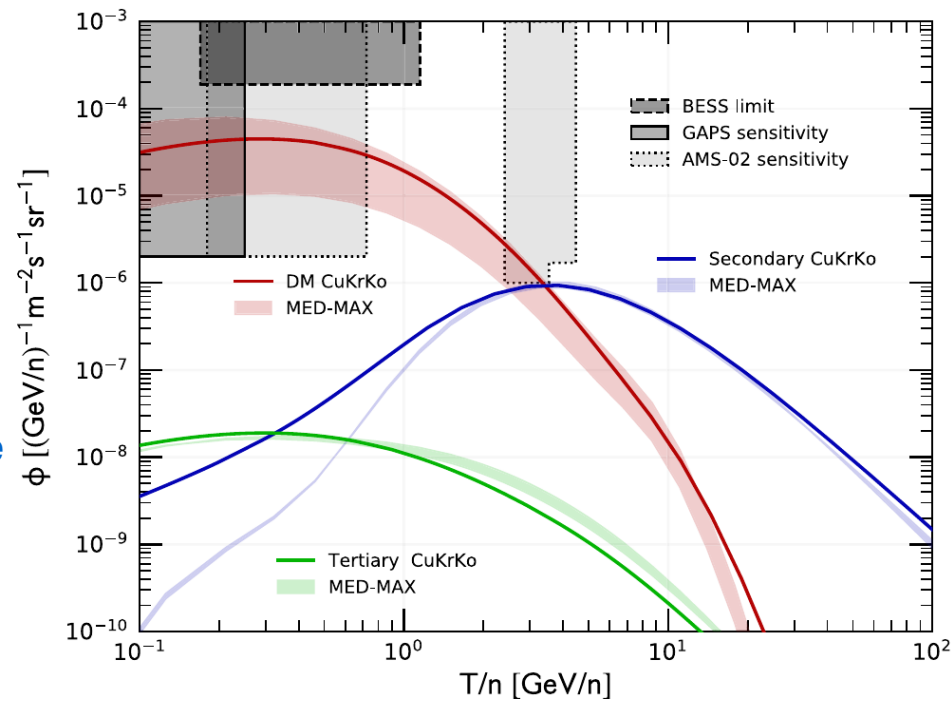
- Versatile collision system
- **Energy scale** $\sqrt{s}_{NN} = 110$ GeV typical of the cosmic ray collisions with the InterStellar Matter
- Accessible range of the anti-proton kinematic spectrum of interest for cosmic processes
- **Crucial measurements** for improving the precision of the **secondary $\bar{p}$ cosmic ray flux** prediction
- Critical to interpret measurement of spaceborn experiments



B₂ measurement improving DM searches



B₂
→
measure



- Predictions are updated with LHC coalescence parameter measurements
- Coalescence momentum p_0 constrained:
- Predictions for DM signal **increased by >10x**
- Predictions for background flux **increased by 2-3x**

$$B_2 = \frac{m_d}{m_p m_n} \frac{\pi p_0^3}{6}$$

$$B_2 \sim 0.01 \rightarrow p_0 \sim 250 \text{ MeV}/c$$

Conclusions and outlook

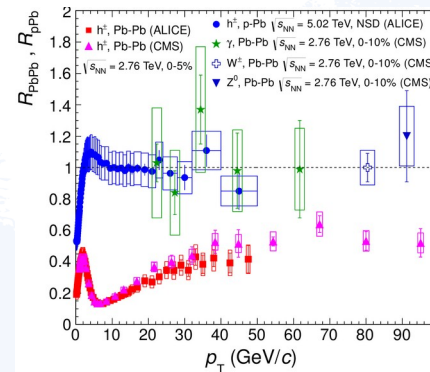
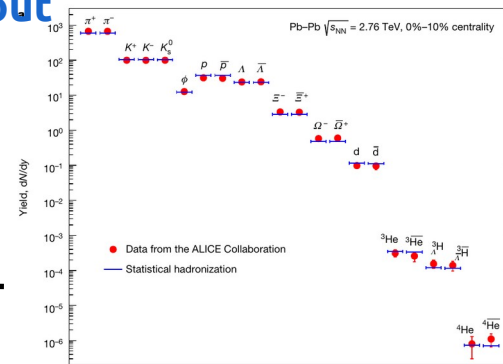
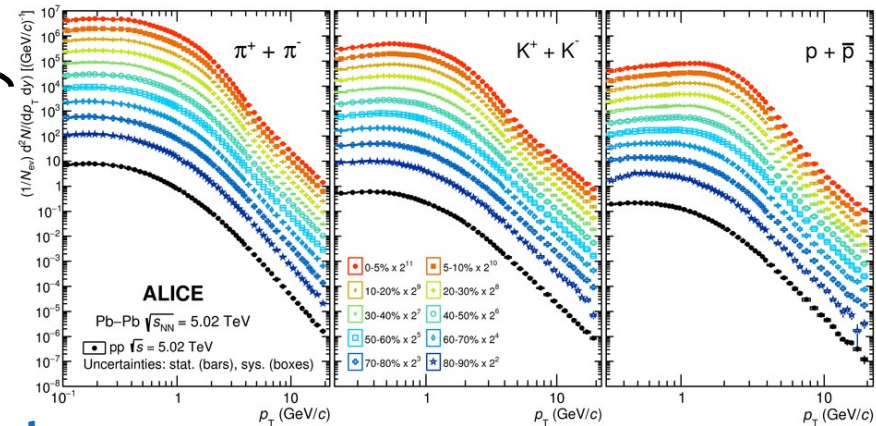
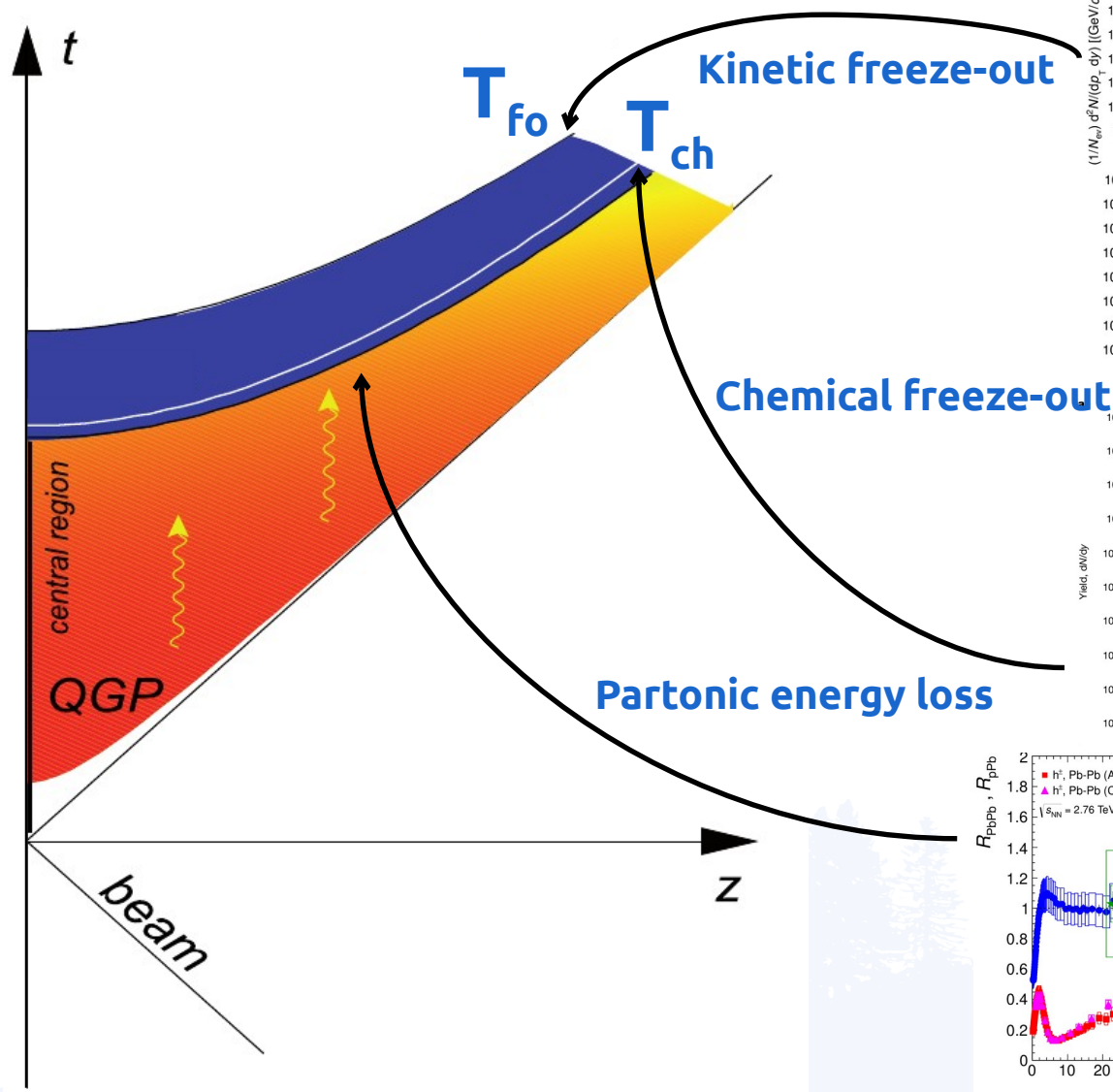


Heavy ion collisions can be used as a large set of "tools"

- Understanding of **soft and hard QCD** processes
 - AA, pA, pp are to be seen as a single block they are all needed to understand the underlying dynamics
- **Antimatter factories** to explore detector effects and contribute to **dark matter searches** → needed from experiments at the collider:
 - Understand antinuclei formation in DM decays
 - Understand antinuclei formation in background reactions
 - Understand interaction of antinuclei with matter to determine the transparency of the galaxy

Future experiments with upgraded detectors will collect data at higher luminosity and providing measurements with **unprecedented precision!**

- Focus will be on **hard probes** (charm and beauty quarks) and on **rare probes** (light nuclei and hypernuclei)
- Impact will be more incisive also on dark matter searches!



And way way more!