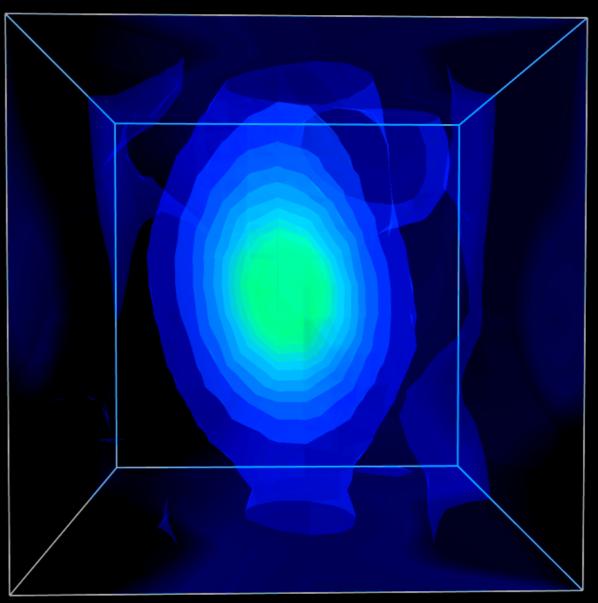
Electroweak Sphaleron in a magnetic field



JA, Kari Rummukainen [arxiv:2301.08626]

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Electroweak Sphaleron in a magnetic field Contents:

- Why is sphaleron in a magnetic field interesting:
 - 1. Sphaleron has a magnetic dipole moment
 - 2. electroweak 'phase transition' gets modified in a magnetic field
- Measuring the sphaleron rate in a magnetic field
- Results

Electroweak Sphaleron

 EW chiral anomaly leads to non-conservation of baryon and lepton number:

$$3\Delta N_{CS} = \Delta B = \Delta L$$

Chern-Simons numbers:

$$N_{\rm CS}(t) \equiv N_{\rm CS}^W(t) - N_{\rm CS}^Y(t)$$

• U(1):
$$N_{\rm CS}^Y(t) \equiv \frac{g'^2}{32\pi^2} \int_0^t {\rm d}t \int {\rm d}^3 x \epsilon_{\alpha\beta\gamma\delta} B^{\alpha\beta} B^{\gamma\delta}$$

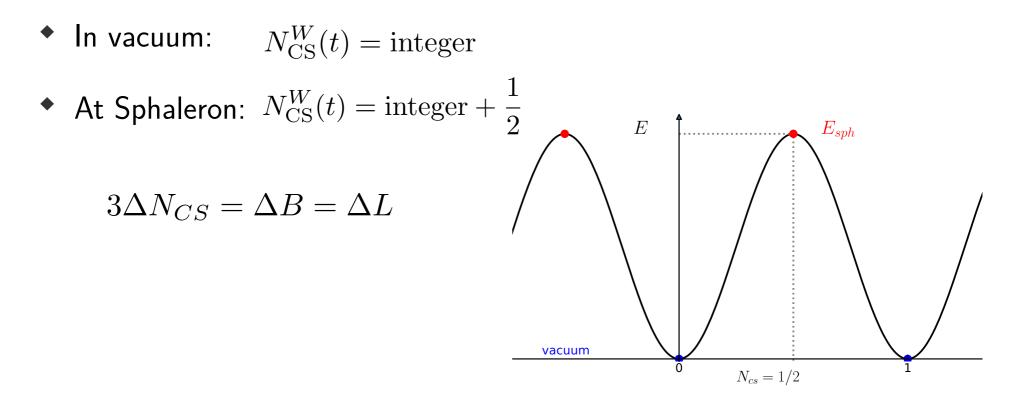
Trivial in vacuum. Identical to zero.

• SU(2):
$$N_{\rm CS}^W(t) \equiv \frac{g^2}{32\pi^2} \int_0^t {\rm d}t \int {\rm d}^3 x \epsilon_{\alpha\beta\gamma\delta} F^{\alpha\beta} F^{\gamma\delta}$$

Non-trivial.

Sphaleron arises from non-trivial topology of SU(2)

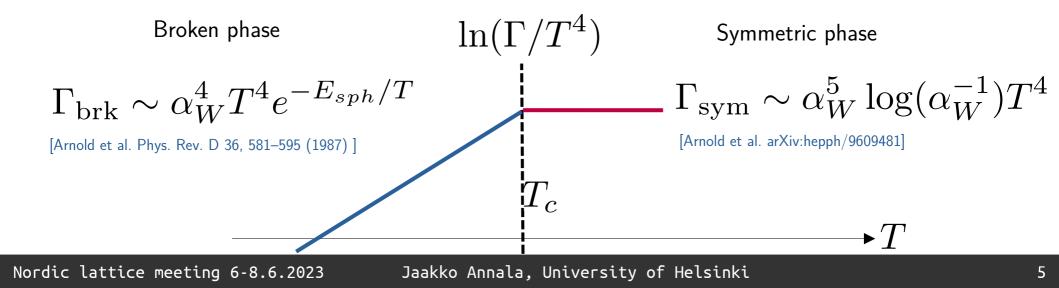
- Infinitely many classically equivalent but topologically different vacua.
- Sphaleron: finite energy solution of classical EoMs separating two topologically distinct vacua.



Sphaleron rate

$$\Gamma = \lim_{V,t \to \infty} \frac{\langle N_{\rm CS}(t)^2 \rangle}{Vt}$$

- How fast sphaleron transitions are happening
- How the rate behaves through the transition is important for Baryo/Lepto-genesis
- Studied extensively in the past without U(1):[D'Onofrio et al. ArXiv:1404.3565, others...]



Why is sphaleron in a magnetic field interesting?

 O. Magnetic fields could have been around during the electroweak phase transition.

1. Electroweak phase transition gets modified

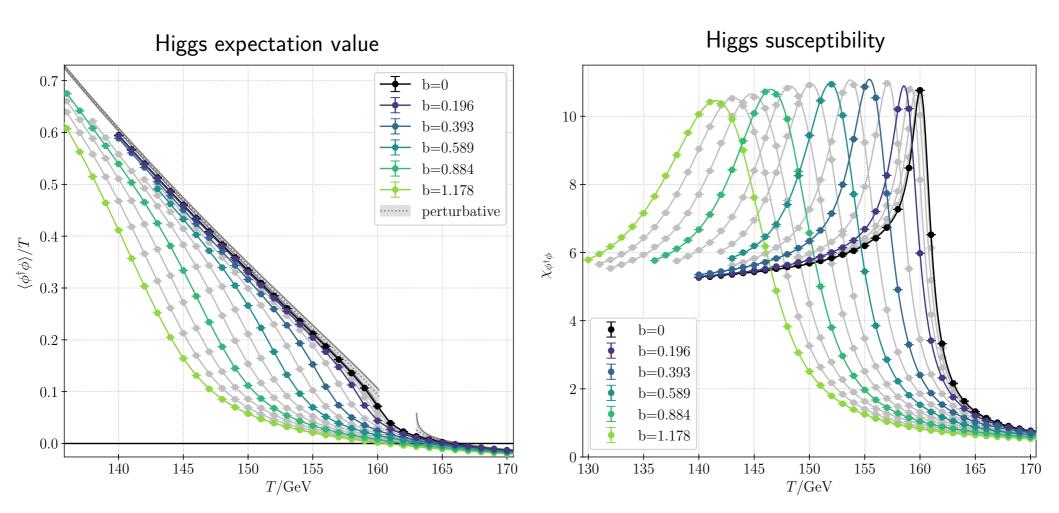
- The details of the electroweak phase transition is changed by an external (hyper)magnetic field.
 - Crossover temperature is shifted to lower temperatures and the "strength" is modified.
 - Higgs expectation value lowered thus sphaleron rate is increased.

- At zero temperature large magnetic fields yield interesting phenomena:
 - Periodic lattice of vortices: Ambjørn-Olesen phase.
- Have not been seen in finite T
 [Kajantie et al. , hep-lat/9809004]

[Ambjørn & Olesen, Phys. Lett. B214, 565–569 (1988)] [Chernodub et al. , 2206.14008]

EW transition shifts to lower temperature

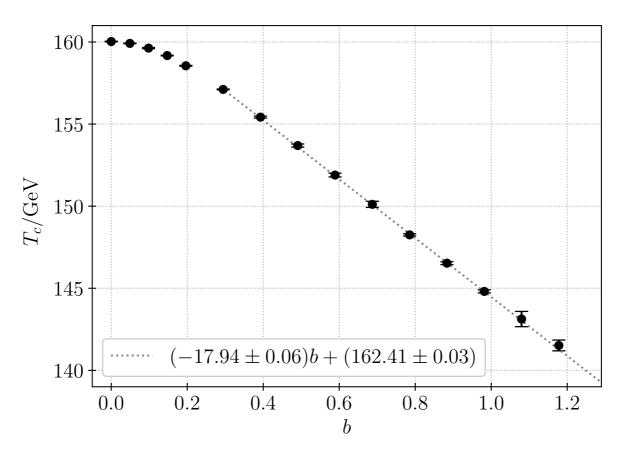
• With increased magnetic field the EW transition shifts to lower temperatures and the transition gets "wider". $B_V^{4d} \simeq 2bT^2$



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EW transition shifts to lower temperature

- With increased magnetic field the EW transition shifts to lower temperatures and the transition gets "wider".
- $T_c = peak$ of the Higgs susceptibility



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 $\frac{g}{a \pi v} = 1.5$

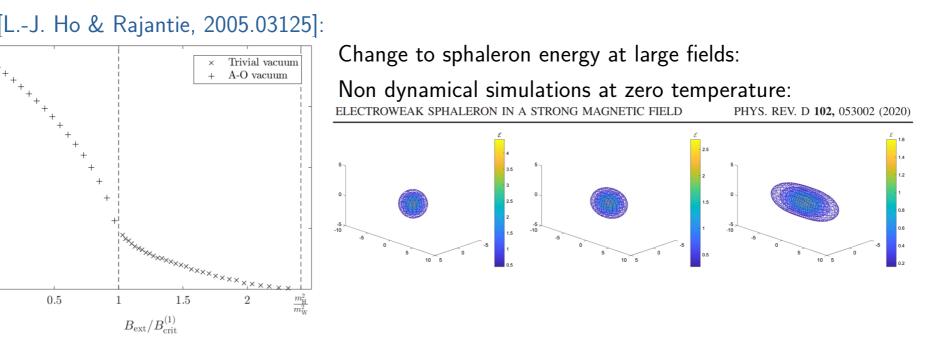
Sphaleron energy

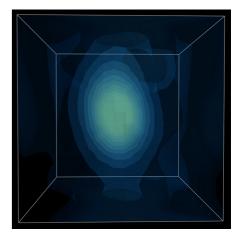
0.5

0

2. Sphaleron has a magnetic dipole moment

- In a magnetic field the energy of the sphaleron can be lowered.
 - For small magnetic fields a dipole interaction is expected: $\Delta E_{\rm sph} = -B_{\rm ext} \cdot \mu_{\rm sph}$
- Sphaleron gets elongated along the magnetic field.





3. U(1) CS number with external magnetic field

- With non-zero magnetic field the U(1) CS number also diffuses.
- Not restricted to integers in vacuum like the SU(2) counter part.
 - Can have any value.
- Can lead to baryon and lepton number change on its own.

[Figueroa et. al., 1707.09967]...

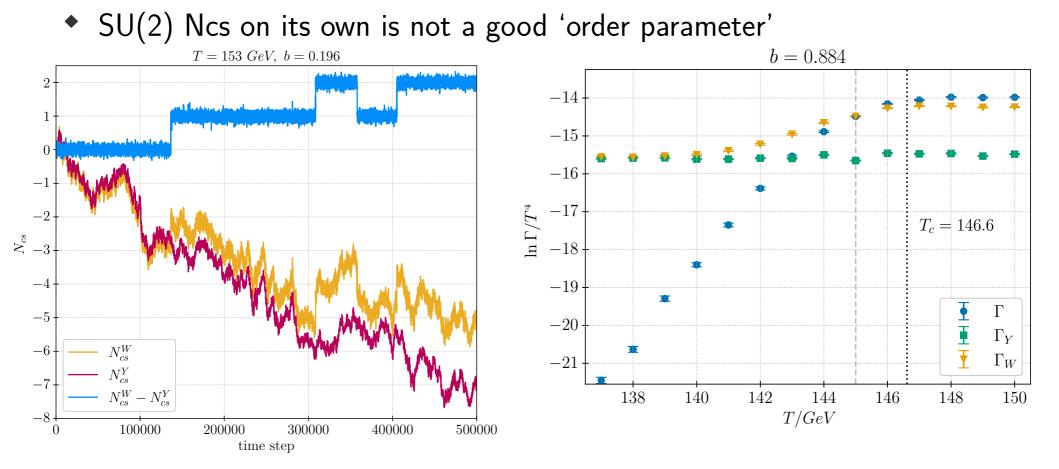
$$N_{\rm CS}(t) \equiv N_{\rm CS}^W(t) - N_{\rm CS}^Y(t),$$

$$3\Delta N_{CS} = \Delta B = \Delta L,$$

$$\Gamma = \lim_{V,t \to \infty} \frac{\langle N_{\rm CS}(t)^2 \rangle}{Vt}$$

In the broken phase SU(2) and U(1) N_{CS} are not independent

 In the broken SU(2) and U(1) CS numbers are highly correlated, only the physical difference of the two gets suppressed, with non-zero magnetic field present.



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Dynamical lattice simulations of effective theory

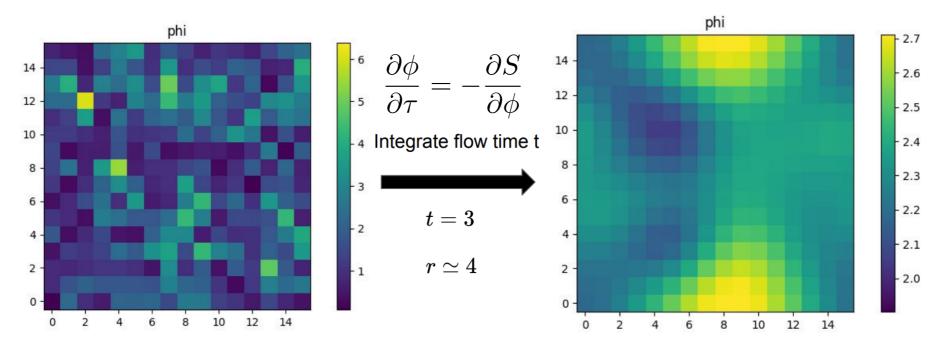
- Effective 3d theory: in finite T get hierarcy of scales $\pi T, gT, g^2T$
 - Temporal and fermionic fields integrated out and we are left with
 3d bosonic theory. [Kajantie et. al., hep-ph/9508379]

$$L = \frac{1}{4}F_{ij}F_{ij} + \frac{1}{4}B_{ij}B_{ij} + (D_i\phi)^{\dagger}(D_i\phi) + m_3^2\phi^{\dagger}\phi + \lambda_3(\phi^{\dagger}\phi)^2$$

- Realtime: dynamics of SU(2) soft modes are described by Langevin equation. (In practice use heatbath update.) [Bödeker, arXiv:hep-ph/9801430.]
 - not true for U(1), correct way to describe the its time evolution is not clear. However, in the broken phase should not affect sphaleron rate.

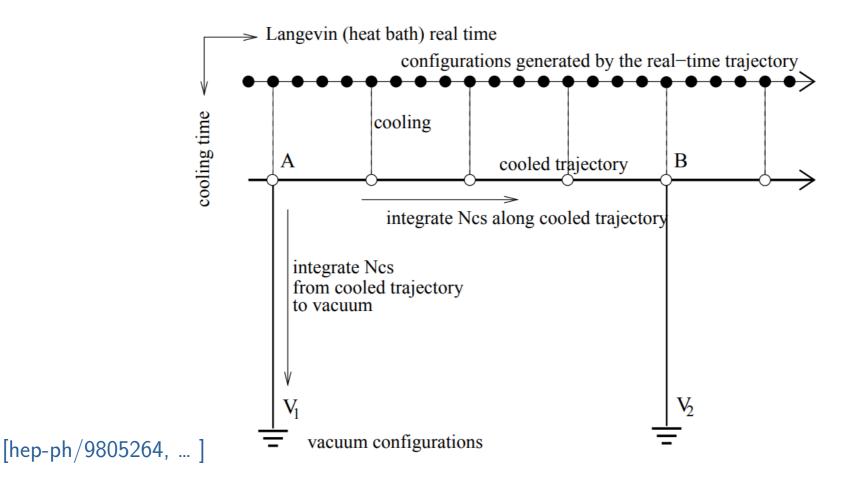
Measuring the sphaleron rate

- Lattice definition of Ncs contains UV noise not related to sphaleron transitions.
- Sphaleron is big in lattice units $~~\sim 1/(g^2T)$
- After every time step smooth the fields using gradient flow



Measuring the sphaleron rate

- Track the Ncs from the cooled fields.
- With a magnetic field we also need to cool the U(1) fields and compute its Ncs.



Multicanonical method

- Sample with probability: $P_W \sim e^{-S + W(N^W_{cs}, N^Y_{cs})}$
 - Weight function W chosen to favor sphalerons
- Run simulation sampling with the weighted distribution to get canonical distribution $P = P_W e^{-W}$ [D'Onofrio et. al., 1207.0685, refs therein...]

$$\Gamma = \frac{P(|N_{cs} - \frac{1}{2}| < \epsilon/2)}{\epsilon V} \langle \left| \frac{\Delta N_{cs}}{\Delta t} \right| \rangle d$$
statistical dynamical

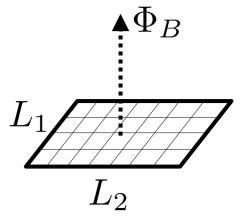
Multicanonical method becomes very inefficient with large volumes

✤ Ncs is a global quantity: increasing lattice size does not help with statistics.

External hypermagnetic field on the lattice

- Maintaining translational invariance on a lattice with a periodic boundary conditions requires to quantize the external magnetic field into flux quanta: $g'_3 \Phi_B = 4\pi n_b$, $n_b \in \mathbb{Z}$
- Magnetic flux density:

$$b \equiv \frac{g_3' B_Y^{3d}}{g_3^4} = \frac{4\pi n_b}{L_1 L_2} \left(\frac{1}{g_3^2 a}\right)^2$$

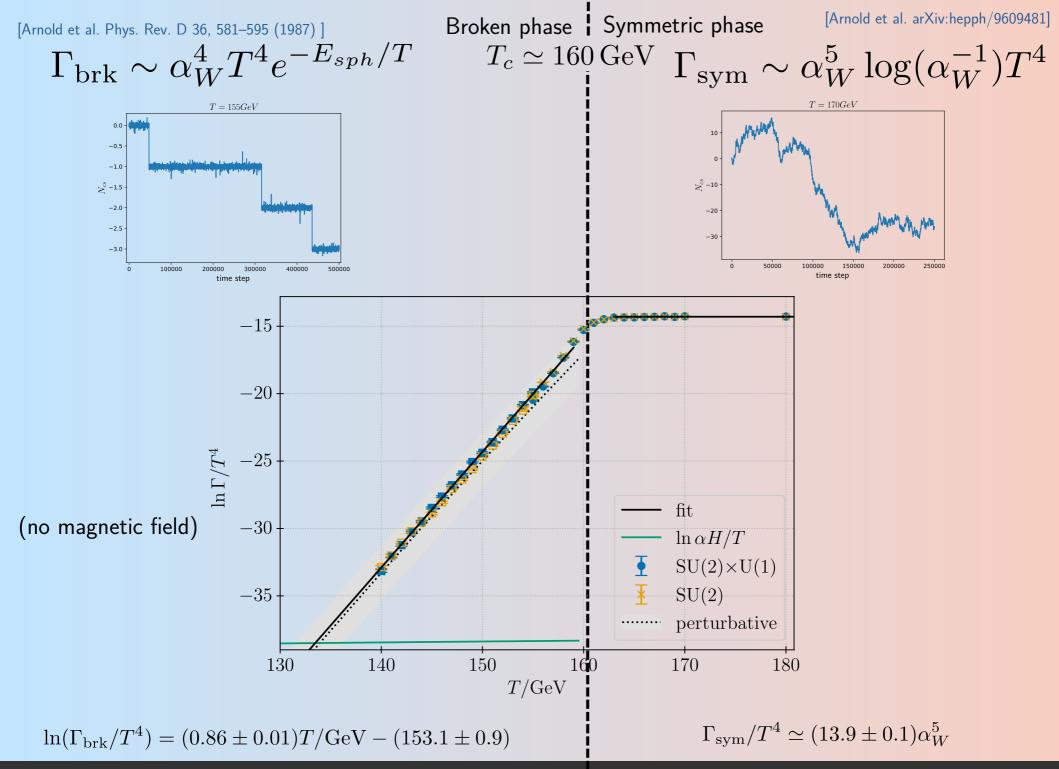


$$B_Y^{4d} \simeq (g'/g^4)bT^2 \approx 2bT^2$$

Volume & Lattice spacing dependence

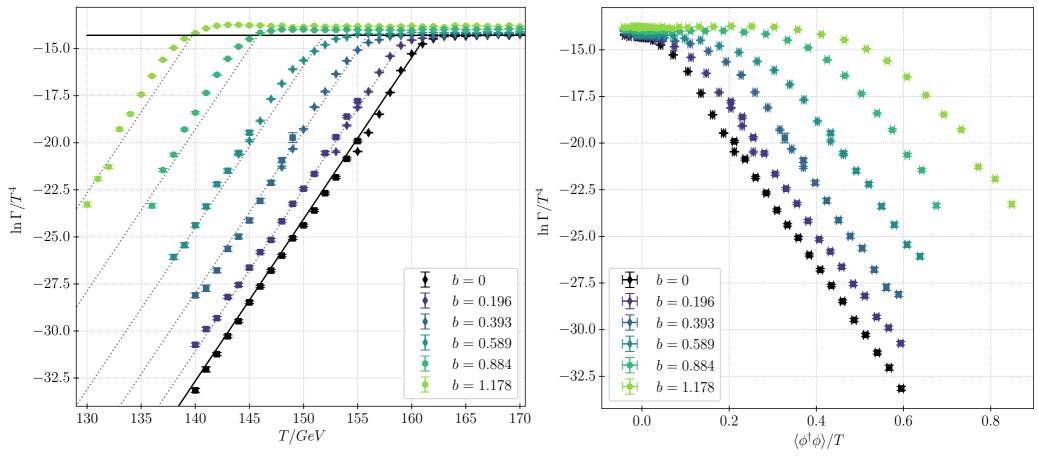
Continuum limit is computationally Volume dependence very mild after $L = 8/g_3^2$ very expensive SU(2), a $g^2T = 0.240$ -16.151.2 ł Ŧ -16.20¥ -16.25 -Ŧ $\Gamma_{sph}^{}/\left(\alpha T\right) ^{4}$ 0.8 T = 157 GeV-16.30 $\ln \Gamma/T^4$ -16.35-22.250.4 ¥ -22.50Ţ₹ -22.750.0 0 10 5 15 T = 150 GeV-23.00 $Lg^{2}T$ [Laine et al., 2209.13804] 0.025 0.0750.1250.000 0.050 0.100 $0.150 \quad 0.175$ $ag_{3}^{2}/4$ [Moore et al. 1011.1167] [Moore&Rummukainen, hep-ph/9906259.]

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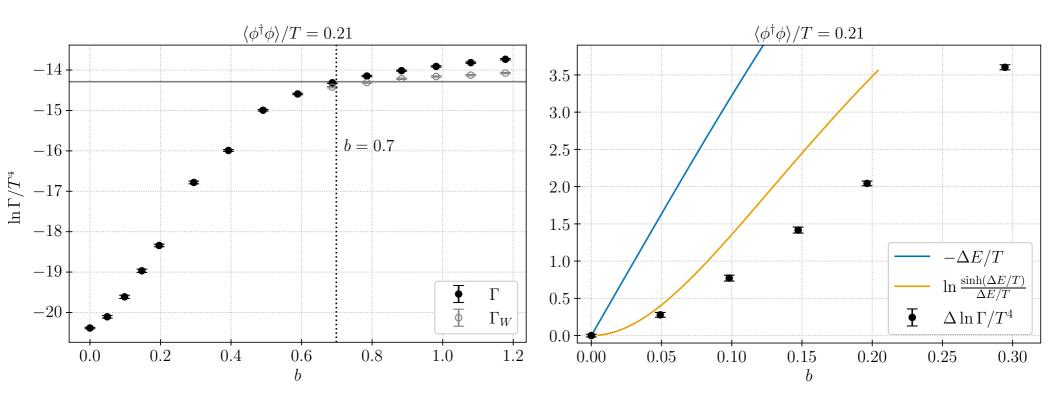
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Sphaleron rate in a magnetic field



- Black solid lines: b=0 fits
- Grey dotted: b=0 fit shifted according to the shift of the pseudo critical temperature.
- For small external fields the sphaleron dipole moment has bigger effect compared to the changing Higgs expectation value.

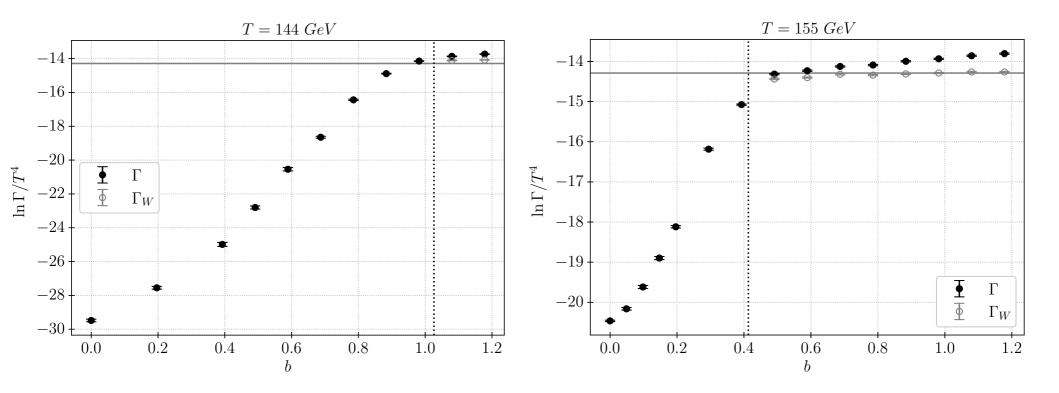
Rate vs magnetic field at constant vev



Solid line: semi analytical computation assuming dipole&spherical approximations

Rate vs magnetic field at constant T

- In the symmetric phase the full rate is slightly faster than its components alone $N_{
 m CS}^W N_{
 m CS}^Y$
- In symmetric phase SU(2) rate does not change significantly.



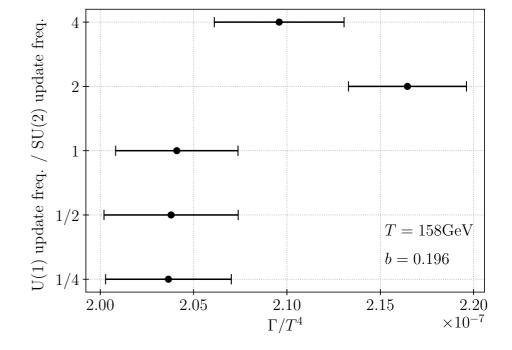
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Conclusions

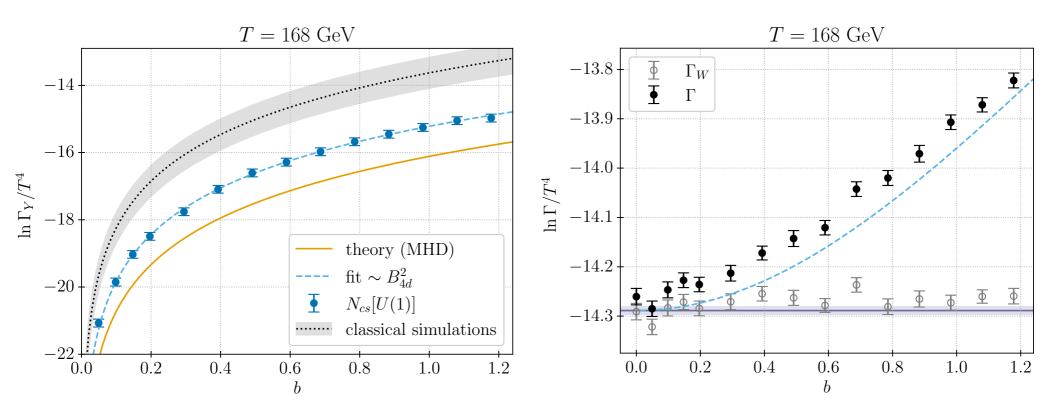
- We performed first dynamical lattice simulations investigating effects of external magnetic field on the sphaleron rate.
- Verified that U(1) does not change the result when magnetic field is zero.
- Sphaleron has a magnetic dipole moment and its energy can be lowered in a magnetic field.
 - For small fields the dipole moment gives the biggest effect. Ultimately the shifting of the pseudocritical temperature dominates.
- Electroweak transition shifts to lower temperatures with increased external magnetic field B_Y = 0...2T²: T_c = 160...145 GeV
- Comparison to semianalytical result: simple dipole approximation works only for very small fields and quickly becomes invalid.

Backup

- Sphaleron size $\sim 1/(g^2T)$
- U(1) modes with wavelength of sphaleron size has time scale also $\sim 1/(g^2T)$ and are weakly coupled $g'^2 \ll g^2$
- * SU(2) is over damped, time scale $\sim 1/(g^2T)^2$



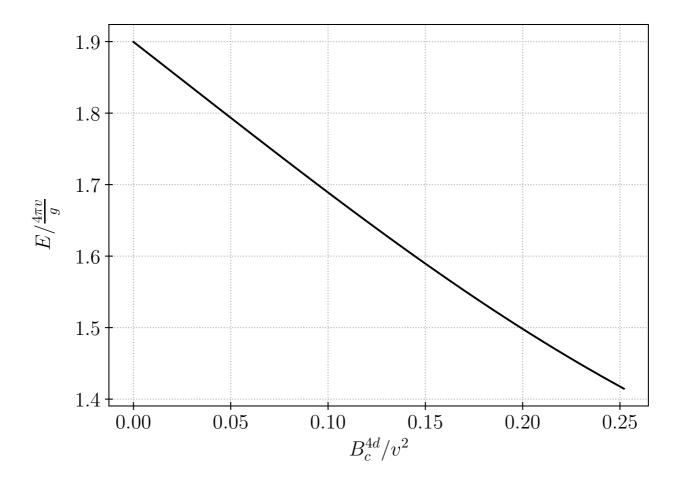
The U(1) rate



(MHD), classical simulations [Daniel G. Figueroa et al. arxiv:1707.09967, arxiv:1904.11892]

Backup

$$\Delta \ln \Gamma / T^4 \sim \ln \left\{ \int \frac{\mathrm{d}\Omega}{4\pi} \exp \left[-\frac{\mu_{\rm sph} B_c^{4d}}{T} \cos \theta \right] \right\} \simeq \ln \frac{\sinh(\Delta E/T)}{\Delta E/T}$$



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