# **GAMBIT-light**

#### Anders Kvellestad

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- 1. What is GAMBIT? (Once more, with feeling)
- 2. Quick GAMBIT news
- 3. GAMBIT-light: What? Why?
- 4. Particles, pandemics and parameter spaces

# 1. What is GAMBIT?









How to test your model against data? The **likelihood** is key!



- Explore the model parameter space ( $\theta_1, \theta_2, \theta_3, ...$ )
- At every point  $\theta$ : compute all predictions( $\theta$ )  $\rightarrow$  evaluate likelihood L( $\theta$ )



 Region of highest L(θ) or lnL(θ): model's best simultaneous fit to all data (but not necessarily a *good* fit, or the most probable θ...)

#### Typical result:

Parameter estimation, presented as profile likelihood and/or posterior density plots



[arxiv:1808.10465]

**Computational challenges:** 

- Need smart exploration of parameter space
- · Need fast theory calculations
- Need fast simulations of experiments (e.g. LHC)
- Need sufficiently detailed likelihoods or full statistical models





<pre>increment signal region counters: / same-sign teprons (preselection 66 nsignalLeptons=2 06 insignalLeptons.at(1)-&gt;pid()&gt;&gt;) { if (signalLeptons.at(0)-&gt;pid()&gt;-signalLeptons.at(1)-&gt;pid()&gt;&gt;) { if (signalLeptons.at(0)abspid()=16 64 signalLeptons.at(0)-&gt;pT()&gt;25)    (signal</pre>
bool pp = false; bool mm = false; iffsignalLeptons.at(0)->pid() > 0)pp = true; iffsignalLeptons.at(0)->pid() < 0)mm = true;
if (num_ISRjets==0) {
// The 0 jet regions if(mT < 100 && pT_ll < 50 && met < 100) _numSR["SS01"]++; if(mT < 100 && pT_ll < 50 && met >= 100 && met < 150 && mmSR["SS02"]++;
if(nT < 100 & 6 pT_ll < 50 & 6 met >= 100 & 6 met < 150 & 6 mm) _numSR["SS03"]++; if(nT < 100 & 6 pT_ll < 50 & 6 met >= 150 & 6 met < 200) _numSR["SS04"]++; if(nT < 100 & 6 pT_ll < 50 & 6 met > 200) _numSR["SS05"]++;
1f(mT < 100 && pT_ll > 50 && met < 100) _numSR["SS06"]++; 1f(mT < 100 && pT_ll > 50 && met >= 100 && met < 150 && pp) _numSR["SS07"]++; 1f(mT < 100 && nT ll > 50 && met >= 100 && met < 150 && mml _numSR["SS08"]++;

Some code infrastructure challenges:

- Need different parameter scanning algorithms
- Need model-agnostic core framework
- · Need to interface *many* external physics codes
- · Need massive parallelisation...
- · ...which implies a need for diskless interfacing
- ...which implies a need to stop external codes from calling STOP and kill your 10,000-CPU scan... :)

# GAMBIT: The Global And Modular BSM Inference Tool

gambit.hepforge.org github.com/GambitBSM EPJC 77 (2017) 784 arXiv:1705.07908

- Extensive model database, beyond SUSY
- Fast definition of new datasets, theories
- Extensive observable/data libraries
- Plug&play scanning/physics/likelihood packages
- Various statistical options (frequentist /Bayesian)
- Fast LHC likelihood calculator
- Massively parallel
- Fully open-source

Members of: ATLAS, Belle-II, CLiC, CMS, CTA, Fermi-LAT, DARWIN, IceCube, LHCb, SHiP, XENON Authors of: BubbleProfiler, Capt'n General, Contur, DarkAges, DarkSUSY, DDCalc, DirectDM, Diver, EasyScanHEP, ExoCLASS, FlexibleSUSY, gamLike, GM2Calc, HEPLike, IsaTools, MARTY, nuLike, PhaseTracer, PolyChord, Rivet, SOFTSUSY, Superlso, SUSY-AI, xsec, Vevacious, WIMPSim



**Recent collaborators**: V Ananyev, P Athron, N Avis-Kozar, C Balázs, A Beniwal, S Bloor, LL Braseth, T Bringmann, A Buckley, J Butterworth, J-E Camargo-Molina, C Chang, M Chrzaszcz, J Conrad, J Cornell, M Danninger, J Edsjö, T Emken, A Fowlie, T Gonzalo, W Handley, J Harz, S Hoof, F Kahlhoefer, A Kvellestad, M Lecroq, P Jackson, D Jacob, C Lin, FN Mahmoudi, G Martinez, H Pacey, MT Prim, T Procter, F Rajec, A Raklev, JJ Renk, R Ruiz, A Scaffidi, P Scott, N Serra, P Stöcker, W. Su, J Van den Abeele, A Vincent, C Weniger, A Woodcock, M White, Y Zhang ++

#### 80+ participants in many experiments and numerous major theory codes

## 2. Quick GAMBIT news



GUT-scale SUSY: 1705.07935



200



1705.07931



stability: 1806.11281



Vector and fermion Higgs portal DM: 1808.10465



GAMET LOG MSSM7

8000

EW-MSSM: 1809.02097



Axion-like particles: 1810.07192



Right-handed neutrinos:



Flavour EFT: 2006.03489



Cosmo ALPs: 2205.13549



More axion-like particles: 2007.05517



Simplified DM, scalar/fermion: 2209.13266



Neutrinos and cosmo: 2009.03287

 $10^{3}$ 

105

ctor DM

 $10^{3}$ 

102





Simplified DM, vector: 2303.08351

Scalar Higgs portal DM w/ vac.



1908.02302





EW-MSSM w/ light gravitino: 2303.09082

- GAMBIT studies since 2017 have required
  >200M CPU hours in total
- Recently got a *EuroHPC Extreme Scale Access* computing grant of **2 x 80M CPU hours**
- On the new LUMI supercomputer (Finland)
- Grant for mapping out pheno implications of combined LHC results and prospects for future runs/colliders





- New web page! gambitbsm.org
- Soon ready #1: GAMBIT + Python scanners
- Soon ready #2: GAMBIT-light



# 3. GAMBIT-light

- GAMBIT-light: GAMBIT without all the physics
- A lightweight yet powerful tool for statistical fits and optimisation tasks
- What GAMBIT-light is not: A full-blown tool for global fits in <your discipline>
   — for that you'd want more of the full GAMBIT functionality
- Key design principles:
  - Users should never need to modify and rebuild any GAMBIT code
  - Minimise the extra maintenance work for GAMBIT developers



Maiken Pedersen (UiO)



Marcin Krotkiewski (UiO)



Brainstorming, early code drafts, testing:

Janina Renk, Fabio Zeiser, Andreas Mjøs, ++

#### • Background:

- We designed GAMBIT to be very general and physics-agnostic
- We put a lot of effort into the main code framework (core framework, config files, scanners, output system, logging, error handling, build system, ...)
- $\circ \rightarrow$  GAMBIT *can* be used for optimisation/fits outside particle/astro/cosmo

#### • Practical experience:

- $\circ$  GAMBIT is a particle physics power tool  $\rightarrow$  fairly heavyweight
- Considerable threshold for non-experts to pick up and use/modify
- In particular: frequent and slow recompilation kills the flow of the early development/experimentation stage of projects

#### • Motivation for GAMBIT-light:

- Help projects outside particle/astro/cosmo use GAMBIT
- Within particle/astro/cosmo: suitable for quick experimentation, MSc projects, "not-so-global fits", etc.





Diver, GreAT, MultiNest, PolyChord, TWalk, grid, random, postprocessor, ...

#### Backends

CaptnGeneral, DarkSUSY, DDCalc, FeynHiggs, FlexibleSUSY, gamLike, gm2calc, HEPLike, HiggsBounds, HiggsSignals, MicrOmegas, nulike, Pythia, SPheno, SUSYHD, SUSYHIT, SuperIso, Vevacious, MontePython, CLASS, AlterBBN, ...



#### 1. Build GAMBIT once

mkdir build cd build cmake -DCMAKE\_BUILD\_TYPE=Release -DWITH\_MPI=On -DCMAKE\_CXX\_COMPILER=g++-11 -DCMAKE\_C\_COMPILER=gcc-11 make -jN scanners # where N is the number of cores to use for the build, e.g. 4 cmake .. # this step is needed for GAMBIT to detect the built scanners make -jN gambit

#### 2. Develop your target/likelihood function code

1	# To import gambit_light_interface, first append the directory containing
2	# gambit_light_interface.so to sys.path. (Alternatively, add this directory
3	# to the PYTHONPATH environment variable.)
4	import sys
5	import os
6	current_dir = os.path.dirname(os.path.abspath(file))
7	<pre>sys.path.append(os.path.join(current_dir, "/lib"))</pre>
8	<pre>import gambit_light_interface as gambit_light</pre>
9	
0	
1	# User-side log-likelihood function, which can be called by GAMBIT-light
2 ~	<pre>def user_loglike(input_names, input_vals, output):</pre>
3	
4	# Make a dictionary of the inputs?
5	input = {input_names[i]: input_vals[i] for i in range(len(input_names))}
6	
7	# Error handling: Report an invalid point using gambit_light.invalid_point.
8	<pre># gambit_light.invalid_point("This input point is no good.")</pre>
9	
Θ	# Error handling: Report a warning using gambit_light.warning.
1	gambit_light.warning("Some warning.")
2	
3	# Error handling: Report an error using gambit_light.error.
4	<pre># gambit_light.error("Some error.")</pre>
5	
6	# Error handling: Error handling, alternative to using gambit_light.error: Throw an exception.
7	<pre># raise Exception("Some exception.")</pre>
8	
9	# Compute loglike
Θ	loglike = input["param_name_1"] + input["param_name_2"] + input["param_name_4"]
1	
2	# Save some extra outputs
3	output["py_user_loglike_output_1"] = 1
4	output["py_user_loglike_output_2"] = 2
5	output["py_user_loglike_output_3"] = 3
6	
7	return loglike
8	

#### 2. Develop your target/likelihood function code

(C++/C/Fortran: build as shared library)

1	#ir	nclude "gambit_light_interface.h"
2		
3	11	User-side log-likelihood function, which can be called by GAMBIT-light.
4	dou	<pre>uble user_loglike(const std::vector<std::string>&amp; input_names,</std::string></pre>
5		<pre>const std::vector<double>&amp; input_vals,</double></pre>
ŝ		std::map <std::string,double>&amp; output)</std::string,double>
7	{	
В		
Э		// Make a map of the inputs?
Э		std::map <std::string,<mark>double&gt; input;</std::string,<mark>
1		<pre>for (size_t i = 0; i &lt; input_names.size(); i++)</pre>
2		{
3		input[input_names[i]] = input_vals[i];
4		}
5		
6		<pre>// Error handling: Report an invalid point using gambit_light_invalid_point.</pre>
7		<pre>// gambit_light_invalid_point("This input point is no good.");</pre>
В		
9		// Error handling: Report a string warning using gambit_light_warning.
Э		gambit_light_warning("Some warning.");
1		
2		<pre>// Error handling: Report an error using gambit_light_error.</pre>
3		<pre>// gambit_light_error("Some error.");</pre>
4		
5		<pre>// Error handling, alternative to using gambit_light_error: Throw a runtime_error.</pre>
6		<pre>// throw std::runtime_error("Some runtime_error.");</pre>
7		
В		// Compute loglike
9		<pre>double loglike = input.at("param_name_2") + input.at("param_name_3");</pre>
Э		
1		// Save some extra outputs
2		output["cpp_user_loglike_output_1"] = 1;
3		output["cpp_user_loglike_output_2"] = 2;
4		<pre>output["cpp_user_loglike_output_3"] = 3;</pre>
5		
6		return loglike;
7	}	1887 9.
в	1	
9	GAN	IBIT LIGHT REGISTER LOGLIKE(user loglike)
Э		0.0119.202.0110.00206.01110.0110.022012.0110.61110.025 0.025 0.025 0.02

#### 3. Configure GAMBIT run with a YAML file

1	UserModel:
2	
3	p1:
4	name: param_name_1
5	prior_type: flat
6	range: [0.0, 5.0]
7	p2:
8	name: param_name_2
9	prior_type: flat
10	range: [0.0, 5.0]
11	p3:
12	name: param_name_3
13	fixed_value: 3.0
14	p4:
15	name: param_name_4
16	<pre>same_as: UserModel::p1</pre>
17	p5-p7:
18	name: param_name_
19	prior_type: flat
20	range: [-1.0, 1.0]
21	
22	UserLogLikes:
23	
24	py_user_loglike:
25	lang: python
26	<pre>user_lib: gambit_light_interface/example_python/example.py</pre>
27	<pre>func_name: user_loglike</pre>
28	output:
29	<pre>- py_user_loglike_output_1</pre>
30	- py_user_loglike_output_2
31	- py_user_loglike_output_3
32	
33	cpp_user_loglike:
34	lang: c++
35	<pre>user_lib: gambit_light_interface/example_cpp/example.so</pre>
36	func_name: user_loglike
37	input:
38	- param_name_2
39	- param_name_3
40	output:
41	- cpp_user_loglike_output_1
42	- cpp_user_loglike_output_2
43	- cpp_user_loglike_output_3
44	

#### 4. Run GAMBIT

mpiexec -np 4 ./gambit -f yaml\_files/your\_configuration\_file.yaml

5. Modify your own code, rerun GAMBIT, modify your own code, rerun GAMBIT, ...



6. Analyse output samples (saved in HDF5 or ascii format)

#### Also: user-supplied prior transformation

#### Python

40	# User-side prior transform function, which can be called by GAMBIT-light.
41	<pre>def user_prior(input_names, input_vals, output):</pre>
42	
43	<pre>for i,v in enumerate(input_vals):</pre>
44	output[i] = v * 10.
45	

#### C++

43	// User-side prior transform function, which can be called by GAMBIT-light.
44	<pre>void user_prior(const std::vector<std::string>&amp; input_names,</std::string></pre>
45	<pre>const std::vector<double>&amp; input_vals,</double></pre>
46	<pre>std::vector<double>&amp; output)</double></pre>
47	{
48	<pre>for (size_t i = 0; i &lt; input_vals.size(); i++)</pre>
49	{
50	<pre>output[i] = input_vals[i] * 10.;</pre>
51	}
52	}
53	
54	GAMBIT_LIGHT_REGISTER_PRIOR(user_prior)

#### Also: user-supplied prior transformation

1	UserModel:
2	
3	p1:
4	name: param_name_1
5	p2:
6	name: param_name_2
7	p3:
8	name: param_name_3
9	p4:
10	name: param_name_4
11	p5-p7:
12	name: param_name_
13	
14	UserPrior:
15	
16	lang: python
17	user_lib: gambit_light_interface/example_python/example_prior_transform.py
18	func_name: user_prior
19	
20	UserLogLikes:
21	
22	py_user_loglike:
23	lang: python
24	<pre>user_lib: gambit_light_interface/example_python/example.py</pre>
25	func_name: user_loglike
26	output:
27	- py_user_loglike_output_1
28	- py_user_loglike_output_2
29	- py_user_loglike_output_3
30	
31	cpp_user_loglike:
32	lang: c++
33	<pre>user_lib: gambit_light_interface/example_cpp/example.so</pre>
34	func_name: user_loglike
35	input:
36	- param_name_2
37	- param_name_3
38	output:
39	<pre>- cpp_user_loglike_output_1</pre>
40	- cpp_user_loglike_output_2
41	- cpp_user_loglike_output_3
42	

# 4. Particles, pandemics and parameter spaces



#### Modelling outbreaks of infectious diseases

- Model with many free parameters
- Non-deterministic dynamics (people)
- Each prediction requires expensive Monte Carlo simulations
- Task: estimate model parameters, make robust predictions

#### **Modelling LHC physics**

- Models with many free parameters
- Non-deterministic dynamics (quantum mechanics)
- Each prediction requires expensive Monte Carlo simulations
- *Task:* estimate model parameters, make robust predictions



- Collaboration with FHI (Norwegian Institute of Public Health)
- Overall goal: Tackle computational challenges common to both particle physics and disease modelling
- Starting point:
  Connect FHI's individual-based model to GAMBIT-light
  → model optimisation and uncertainty estimation
- Some possible next steps:
  - Improved multi-level parallelisation schemes (GPUs?)
  - Develop + use fast ML-based surrogate models (e.g. with *continual learning*)
  - Tailored parameter sampling algorithms
  - o ...

- Ida-Marie (UiO + FHI)
- Are (UiO)
- Me (UiO)
- Jørgen Eriksson Midtbø (FHI)
- Birgitte Freiesleben de Blasio (FHI)
- Francesco Di Ruscio (FHI)
- Yat Hin Chan (FHI)